



Scaling relations for heat and momentum transport in sheared Rayleigh-Bénard convection

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We provide scaling relations for the Nusselt number Nu and the friction coefficient C_S in sheared Rayleigh–Bénard convection, i.e. in Rayleigh–Bénard flow with Couette- or Poiseuille-type shear forcing, by extending the Grossmann & Lohse (*J. Fluid Mech.*, vol. 407, 2000, pp. 27–56, *Phys. Rev. Lett.*, vol. 86, 2001, pp. 3316–3319, *Phys. Rev. E*, vol. 66, 2002, 016305, *Phys. Fluids*, vol. 16, 2004, pp. 4462–4472) theory to sheared thermal convection. The control parameters for these systems are the Rayleigh number *Ra*, the Prandtl number *Pr* and the Reynolds number *Res* that characterises the strength of the imposed shear. By direct numerical simulations and theoretical considerations, we show that, in turbulent Rayleigh–Bénard convection, the friction coefficients associated with the applied shear and the shear generated by the large-scale convection rolls are both well described by Prandtl's (*Ergeb. Aerodyn. Vers. Gött.*, vol. 4, 1932, pp. 18–29) logarithmic friction law, suggesting some kind of universality between purely shear-driven flows and thermal convection. These scaling relations hold well for $10^6 \leq Ra \leq 10^8$, $0.5 \leq Pr \leq 5.0$, and $0 \leq Res \leq 10^4$.

Key words: turbulence simulation, turbulent convection

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Figure 1. Schematic of the (a) CRB and (b) PRB systems.

1. Introduction

The interplay between buoyancy and shear in mixed thermal convection can be studied by either adding Couette-type forcing to the Rayleigh–Bénard (RB) system (Ahlers, Grossmann & Lohse 2009; Lohse & Xia 2010; Chilla & Schumacher 2012; Xia 2013; Shishkina 2021; Ahlers *et al.* 2022; Lohse & Shishkina 2023) to obtain the Couette–RB (CRB) system (Deardorff 1965; Ingersoll 1966; Hathaway & Somerville 1986; Domaradzki & Metcalfe 1988; Solomon & Gollub 1990; Shevkar *et al.* 2019; Blass *et al.* 2020, 2021), or by applying a Poiseuille-type forcing to obtain the Poiseuille–RB (PRB) system (Scagliarini, Gylfason & Toschi 2014; Zonta & Soldati 2014; Scagliarini *et al.* 2015; Pirozzoli *et al.* 2017). A schematic of the two systems is shown in figure 1.

The CRB and PRB systems are described by the incompressible Navier–Stokes equations, the continuity equation and the temperature transport equation, within the Boussinesq approximation. In Cartesian coordinates, they read

$$\partial_t u_i + u_j \partial_j u_i = -\rho^{-1} \partial_i p + \nu \partial_j^2 u_i + \beta g \delta_{i3} \theta + \Pi \delta_{i1}, \quad \partial_i u_i = 0, \tag{1.1a,b}$$

$$\partial_t \theta + u_j \partial_j \theta = \kappa \partial_j^2 \theta, \qquad (1.2)$$

where $\mathbf{u} \equiv (u_x, u_y, u_z)$ is the velocity, p the pressure, θ the reduced temperature, ρ the density of the fluid, g the acceleration due to gravity antiparallel to the z direction, β the isobaric thermal expansion coefficient, v the kinematic viscosity, κ the thermal diffusivity and H is the distance between the horizontal walls. At the top wall (z = H), the reduced temperature is set to $\theta = -\Delta/2$ while at the bottom wall (z = 0), the reduced temperature is set to $\theta = \Delta/2$. For the CRB system $\Pi = 0$, the bottom wall is at rest and a velocity of $2U_w$ is imposed on the top wall. For the PRB system, no-slip conditions are enforced at the walls and a volume forcing Π is applied in the streamwise direction such that it induces a bulk velocity of U_b averaged over the domain volume and time (for the details of the implementation of the shear forcing in the numerical simulations, we refer the reader to § 3). The streamwise direction is oriented along x and the spanwise direction along y. The aspect ratios of the system are defined by $\Gamma_x = L_x/H$ and $\Gamma_y = L_y/H$, with L_x , L_y being the dimensions of the system in the x and y directions, respectively.

The control parameters for the systems are the Rayleigh number, the Prandtl number and the Reynolds number associated with the shear forcing

$$Ra \equiv \frac{\beta g H^3 \Delta}{\nu \kappa}, \quad Pr \equiv \frac{\nu}{\kappa}, \quad Re_S \equiv \frac{U_S H}{\nu}.$$
 (1.3*a*-*c*)

The characteristic velocity scale associated with the shear forcing U_S is given by $U_S \equiv U_w$ for the CRB system and $U_S \equiv U_b$ for the PRB system. Although U_b is formally a response parameter, in our numerical simulations the volume forcing term Π is computed at each

time step to ensure a constant mass flow rate (Quadrio, Frohnapfel & Hasegawa 2016) dictated by U_b , making it a control parameter in our case. The shear forcing for the CRB system is given by the wall Reynolds number $Re_w \equiv U_w H/v$ whereas the shear forcing for the PRB system is given by the bulk Reynolds number $Re_b \equiv U_b H/v$. Henceforth, we use Re_s to indicate shear forcing in equations that are applicable to both CRB and PRB systems.

Similarly, one can also define a Reynolds number associated with the thermal forcing in these systems. From the input parameters, we can construct a Reynolds number $Re_F = U_F H/\nu \equiv \sqrt{Ra/Pr}$, using the free fall velocity scale $U_F = \sqrt{g\beta\Delta H}$. However, the free fall scale is often not a reliable estimate of the flow velocities that develop in natural convection flows. A more appropriate approach is to define the Reynolds number associated with the large-scale convection (LSC) roll given by $Re_L \equiv U_L H/\nu$, with U_L indicating the mean velocity of the 'wind of turbulence' generated by the LSC roll. The parameter Re_L is, however, a response parameter whose variation with Ra, Pr and Re_S is not known a priori.

In the limiting case of zero shear forcing, we can distinguish the Reynolds number associated with the LSC roll in pure RB flow as $Re_R(Ra, Pr) \equiv Re_L(Ra, Pr, Re_S = 0)$. The dynamics of the sheared RB systems are governed by a combined effect of both shear and thermal forcing. Therefore, we can also introduce the Reynolds number Re_T , which is constructed using the total velocity U_T comprising a vector sum of U_S and U_L . Naturally, the time-averaged wall shear stress τ_T generated by the total velocity U_T is also determined by the combined effect of τ_L , which is the time-averaged shear stress locally generated on the walls by the LSC roll, and τ_S , which is the mean streamwise shear stress generated on the walls due to the applied shear forcing. These shear stresses can be expressed in dimensionless form using the friction coefficients associated with the total shear, the LSC roll and the streamwise shear respectively, as

$$C_T \equiv \frac{2\tau_T}{\rho U_T^2}, \quad C_L \equiv \frac{2\tau_L}{\rho U_I^2}, \quad C_S \equiv \frac{2\tau_S}{\rho U_S^2}. \tag{1.4a-c}$$

Once again, in the limiting case of zero shear forcing, we can distinguish the friction coefficient associated with the LSC roll in pure RB flow as $C_R(Ra, Pr) \equiv C_L(Ra, Pr, Re_S = 0)$. The non-dimensional heat flux from the hot bottom wall to the cold top wall is given by the Nusselt number

$$Nu \equiv \frac{\langle u_z \theta - \kappa \, \partial_z \theta \rangle_{A,t}}{\kappa \, \Delta H^{-1}},\tag{1.5}$$

with $\langle \rangle_{A,t}$ indicating the averaging in time and over any horizontal plane spanned by *x* and *y*. A summary of all response and control parameters discussed above is given in table 1 for reference. In this study, we are primarily interested in understanding the dependence of the response parameters on the control parameters, and the physics underlying the connections between the response parameters.

Using the exact relations for the global kinetic and thermal dissipation rates, Grossmann & Lohse (2000, 2001, 2002, 2004) offered a unifying theory for RB convection (hereafter referred to as the GL-theory). For cylinders of unit aspect ratio, Stevens *et al.* (2013) have demonstrated that fitting the GL-theory at four data points from Funfschilling *et al.* (2005) at $Ra = 2.96 \times 10^7$ and $Ra = 1.92 \times 10^{10}$ with Pr = 4.38, from Xia, Lam & Zhou (2002) at $Ra = 2.24 \times 10^8$ with Pr = 554 and from Kerr & Herring (2000) at $Ra = 10^7$ with Pr = 0.07 using four free parameters can predict *Nu* within 4% of experimental and numerical results in most of the parameter space given by $10^4 \le Ra \le 10^{14}$ and $10^{-3} \le 10^{14}$.

Parameter	Definition	Туре	Short description
Ra	$(\beta g H^3 \Delta)/(\nu \kappa)$	Control	Rayleigh number
Pr	ν/κ	Control	Prandtl number
Re_F	$U_F H/v \equiv \sqrt{Ra/Pr}$	Control	Reynolds number associated with free-fall velocity
Re_w	$U_w H/v$	Control	Wall Reynolds number for CRB system
Re_b	$U_b H/v$	Control	Bulk Reynolds number for PRB system
Res	U_SH/v	Control	Reynolds number associated with shear forcing
ReL	$U_L H/ u$	Response	Reynolds number associated with LSC rolls for sheared RB
<i>Re_R</i>	$U_R H/\nu \equiv Re_L (Re_S = 0)$	Response	Reynolds number associated with LSC rolls for pure RB ('wind Reynolds number')
Re _T	$U_T H/v$	Response	Reynolds number computed using the total velocity
C_S	$2\tau_S/(ho U_S^2)$	Response	Friction coefficient associated with shear forcing
C_L	$2 au_L/(ho U_L^2)$	Response	Friction coefficient associated with LSC rolls for sheared RB
C_R	$2\tau_R/(\rho U_R^2) \equiv C_L(Re_S = 0)$	Response	Friction coefficient associated with LSC rolls for pure RB
C_T	$2 au_T/(ho U_T^2)$	Response	Friction coefficient computed using the total velocity
Nu	$\langle u_z \theta - \kappa \partial_z \theta \rangle_{A,t} / (\kappa \Delta H^{-1})$	Response	Nusselt number
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Table 1. A summary of all control and response parameters discussed in § 1.

 $Pr \le 10^2$, with only two small ranges that exhibit a greater than 10 % disagreement. In the same paper Stevens *et al.* (2013), a fit of similar quality is achieved also for an aspect ratio $\Gamma = 1/2$, with slightly different prefactors. This work has further been extended by Ahlers *et al.* (2022) to include the effects of aspect ratios between 1/32 and 32. The GL-theory has also been extended to the ultimate regime (Grossmann & Lohse 2011), where the heat transport is considerably enhanced, as the laminar-type boundary layers become turbulent due to a non-normal–nonlinear instability; for a detailed discussion see Roche (2020) and Lohse & Shishkina (2023).

Presently, the GL-theory has been applied to RB convection without imposed shear. The objective of this work is to extend the theoretical approach to sheared thermal convection. Scagliarini *et al.* (2014, 2015); Pirozzoli *et al.* (2017); Blass *et al.* (2020, 2021) have made progress in understanding the variation of heat transfer in RB convection with imposed shear. Scagliarini *et al.* (2014, 2015) proposed a model based on the concept of eddy viscosity and eddy diffusivity to explain the counter-intuitive initial decrease and subsequent increase in Nu with increasing Re_b for the PRB system. Blass *et al.* (2020) observed a similar effect in the CRB system and attributed it to the initial destruction of the large-scale flow organisation and the subsequent formation of large meandering flow structures (Hutchins & Marusic 2007). They divided the flow into a buoyancy-dominated, a transitional and a shear-dominated regime, based on the Monin–Obukhov length scale. Blass *et al.* (2021) further investigated the effect of Pr on the variation of Nu with Re_w and concluded that the non-monotonic behaviour of $Nu(Re_w)$ is a consequence of flow

layering, plume sweeping and bulk heat entrapment. Building on these findings, in this paper we propose a more general formulation applicable to sheared thermal convection, in the spirit of the GL-theory.

The objective of this paper is to extend the GL-theory to CRB and PRB systems by providing scaling relations for $Nu(Ra, Pr, Re_S)$ and $C_S(Ra, Pr, Re_S)$. Furthermore, we will show similarities in the dependence of C_S on Re_S in Couette and Poiseuille flows and C_L on Re_L in RB convection, suggesting some sort of universality in the behaviour of the friction coefficient in shear-driven flows and thermal convection. In § 2, we build the theoretical framework, using rigorous relations for globally averaged kinetic and thermal dissipation rates. In § 3, we validate the theoretical scaling relations against direct numerical simulations (DNS). Finally, the conclusions are presented in § 4.

2. Extending the Grossmann-Lohse theory to CRB and PRB

2.1. Kinetic and thermal dissipation rates

To extend the GL-theory to sheared thermal convection, we formulate exact global relations for the kinetic (ϵ_u) and thermal (ϵ_θ) dissipation rates in the CRB and PRB systems. These arise from the time- and volume-averaged equations for the kinetic energy and temperature variance, respectively. The relation for the mean thermal dissipation rate is the same as in the classical RB convection

$$\epsilon_{\theta} = \left\langle \kappa \left(\partial_{j} \theta \right)^{2} \right\rangle_{V,t} = \kappa \frac{\Delta^{2}}{H^{2}} N u, \qquad (2.1)$$

see e.g. Shraiman & Siggia (1990) and Siggia (1994). The relation for the mean kinetic dissipation rate reads

$$\epsilon_{u} = \left\langle \nu(\partial_{j}u_{i})^{2} \right\rangle_{V,t} = \frac{\nu^{3}}{H^{4}} \left(\underbrace{(Nu-1)RaPr^{-2}}_{\text{Buoyancy term}} + \underbrace{C_{S}Re_{S}^{3}}_{\text{Shear term}} \right), \quad (2.2)$$

with $\langle ... \rangle_{V,t}$ indicating the average over time and volume. Note that the expression for thermal dissipation (2.1) is the same as that in classical RB convection but the expression for kinetic dissipation (2.2) includes contributions from both buoyancy and shear forcing.

2.2. Kinetic energy, large-scale convection rolls and boundary layer thickness

One of the central ideas of the GL-theory is the presence of persistent LSC rolls that churn the bulk of the system and generate boundary layers at the walls. As a result, the mean kinetic energy of the RB system is expected to scale as $\sim U_R^2$, where U_R is the velocity scale of the LSC roll (in the absence of any shear forcing). In pure RB flow, this mean kinetic energy is solely generated by the buoyancy forcing. However, in the case of sheared RB flow, where the LSC roll has a velocity scale U_L , the mean kinetic energy consists of contributions from both the LSC roll and the imposed shear flow. Therefore, we add the kinetic energy of the mean flow U_S^2 and the associated turbulent kinetic energy (TKE), which scales as the square of the friction velocity $u_{\tau} = \sqrt{\tau_S/\rho}$, to write

$$U_T^2 \approx U_L^2 + U_S^2 + 2\gamma u_{\tau}^2,$$
 (2.3)

with γ being a prefactor. This can also be written in terms of the corresponding Reynolds numbers as

$$Re_T^2 \approx Re_L^2 + Re_S^2 + \gamma C_S Re_S^2.$$
(2.4)

For a laminar Prandtl–Blasius-type (Prandtl 1904; Blasius 1908) boundary layer, there is no TKE, so the contribution $C_S Re_S^2$ vanishes and we can approximate (2.4) as

$$Re_T^2 \approx Re_L^2 + Re_S^2. \tag{2.5}$$

An interpretation of the above equation (2.5) is that the velocity associated with the LSC roll is preferentially oriented along the direction orthogonal to the shear forcing, consistent with the flow structures observed by Pirozzoli *et al.* (2017); Blass *et al.* (2020). Note that the validity of (2.5) is limited to sufficiently low values of Re_S wherein the contribution to spanwise shear stresses from shear forcing is negligible in comparison with the contribution from the LSC rolls. At high shear forcing, the spanwise shear stresses generated by velocity fluctuations arising purely from shear forcing may no longer be negligible, in which case (2.5) no longer holds.

In the buoyancy-dominated regime, relation (2.5) can be better understood by considering many LSC rolls each orientated at an angle α with the streamwise direction and studying the probability distribution of α given by $\phi(\alpha)$. Here, we mean LSC rolls in a broad sense, without addressing the exact details of the flow organisation at this stage. Empirical observations regarding flow organisation are reported in section § 3.3. Since the total velocity U_T arises from a vector addition of the shear velocity and the LSC velocities, we can express it in terms of $\phi(\alpha)$ as

$$U_T^2 = \int_0^{2\pi} \phi(\alpha) \left(U_S^2 + U_L^2 + 2U_S U_L \cos(\alpha) \right) d\alpha.$$
 (2.6)

Due to the symmetry of the system and the periodic boundary conditions in the horizontal directions, it is reasonable to consider that there are an equal number of clockwise and counter-clockwise LSC rolls within the sheared RB system. When averaged over the entire volume of the system, we postulate that the probability distribution $\phi(\alpha)$ is symmetric about the spanwise direction, i.e. about $\alpha = \pm \pi/2$. Applying this symmetry condition to (2.6) gives us the relation (2.5). The velocity scale U_T associated with the kinetic energy of the system can thus be considered as a vector sum of perpendicular velocity contributions from the shear forcing in the streamwise direction and the LSC roll in the spanwise direction. Following this approach, we can also decompose the total shear stress τ_T generated by U_T into the streamwise component $\tau_s = \tau_T U_S/U_T$, generated by U_S , and the spanwise component $\tau_L = \tau_T U_L/U_T$, generated by U_L . This assumption gives us three equivalent definitions of the kinetic boundary layer thickness λ_u , namely

$$\lambda_u \equiv \frac{2H}{C_L R e_L} = \frac{2H}{C_S R e_S} = \frac{2H}{C_T R e_T}.$$
(2.7)

Here, we used the slope criterion from Shishkina *et al.* (2010) for the definition of the kinetic boundary layer thickness. Similarly, we define the thermal boundary layer thickness λ_{θ} also with the slope criterion as

$$\lambda_{\theta} \equiv \frac{H}{2Nu}.$$
(2.8)

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2.3. Estimating bulk and boundary layer contributions

Another key idea of the GL-theory is the splitting of ϵ_u and ϵ_{θ} into their bulk and boundary layer contributions as follows:

$$\epsilon_u = \epsilon_{u,BL} + \epsilon_{u,bulk}, \quad \epsilon_\theta = \epsilon_{\theta,BL} + \epsilon_{\theta,bulk}, \quad (2.9a,b)$$

with $\epsilon_{u,BL}$ being the boundary layer contribution to the kinetic dissipation, $\epsilon_{u,bulk}$ being the bulk contribution to the kinetic dissipation, $\epsilon_{\theta,BL}$ being the boundary layer contribution to the thermal dissipation and $\epsilon_{\theta,bulk}$ being the bulk contribution to the thermal dissipation. Therefore, we focus on estimating $\epsilon_{u,BL}$, $\epsilon_{u,bulk}$, $\epsilon_{\theta,BL}$ and $\epsilon_{\theta,bulk}$ for sheared RB systems.

First, we estimate $\epsilon_{u,BL}$ using (2.7) as

$$\epsilon_{u,BL} \sim \nu \frac{U_T^2}{\lambda_u^2} \frac{\lambda_u}{H} \sim \frac{\nu^3}{H^4} C_T R e_T^3 \equiv \frac{\nu^3}{H^4} \left(\underbrace{C_L R e_L^3}_{\text{LSC term}} + \underbrace{C_S R e_S^3}_{\text{Shear term}} \right), \quad (2.10)$$

which is, in turn, a sum of contributions from the LSC rolls and the applied shear forcing. Note that the shear contribution here exactly matches that in the global relation (2.2). Next, we estimate $\epsilon_{u,bulk}$. Here, it is important to note that the bulk dissipation rate is dominated by the contribution from LSC rolls, while the contribution from applied shear forcing is much smaller. Therefore, we only focus on estimating the contribution to $\epsilon_{u,bulk}$ from the LSC rolls by assume that the velocity scale U_L associated with the LSC rolls is responsible for stirring the bulk with a kinetic energy that scales with U_L^2 . However, LSC rolls are swept at the boundary layer height due to the applied shear forcing and the time scale of the stirring process is governed not by the velocity U_L associated with the LSC rolls but by U_T which is the total velocity scale. As we shall find later in § 3.2, this is a key assumption that explains the trend of Nu with increasing Re_S in the buoyancy-dominated regime. Following these assumptions, we write

$$\epsilon_{u,bulk} \sim U_L^2 \frac{U_T}{H} \equiv R e_L^2 R e_T.$$
(2.11)

At high-shear forcing, the contribution of the boundary layer to the dissipation rate arising from shear forcing dominates the bulk contribution from LSC rolls. In the limiting case of passive transport in Couette/Poiseuille flow, one can rigorously derive that the total kinetic dissipation rate $\epsilon_u = (v^3/H^4)C_SRe_S^3$. In this sense, the kinetic dissipation rate of sheared RB at high shear forcing will always be dominated by the boundary layer contribution and estimating $\epsilon_{u,bulk}$ in the shear-dominated regimes is redundant.

The analogous estimate for $\epsilon_{\theta,BL}$

$$\epsilon_{\theta,BL} \sim \kappa \frac{\Delta^2}{\lambda_{\theta}^2} \frac{\lambda_{\theta}}{H} \sim \kappa \frac{\Delta^2}{H^2} N u,$$
(2.12)

is identical to the exact relation (2.1), on the one hand showing consistency of the approach, but on the other hand not giving new information. Therefore, following the GL-theory, we match the magnitude of the advective and diffusive terms of (1.2) at the thermal boundary layer height to obtain

$$u_{\rm v}\partial_{\rm v}\sim\kappa\,\partial_{zz}.$$
 (2.13)

As in the GL-theory, for regimes where the thermal boundary layer is thicker than the kinetic boundary layer ($\lambda_{\theta} > \lambda_{u}$, associated with low *Pr*), we estimate $u_{y} \sim U_{L}$, $\partial_{y} \sim 1/H$,

and $\partial_{zz} \sim \lambda_{\theta}^{-2}$. Using these estimates in (2.13) with λ_u from (2.7) and λ_{θ} from (2.8), we obtain

$$Nu \sim Pr^{1/2}Re_L^{1/2} \equiv \frac{Pr^{1/2}}{C_L Re_L^{1/2}}C_T Re_T.$$
 (2.14)

For high Pr regimes, where $\lambda_{\theta} < \lambda_{u}$, we estimate $u_{y} \sim U_{L}\lambda_{\theta}/\lambda_{u}$ due to the fact that the relevant velocity scale at the thermal boundary height is smaller than the velocity U_{L} by a factor $\lambda_{\theta}/\lambda_{u}$, exactly as in the GL-theory. Once again using the estimates $\partial_{y} \sim 1/H$, and $\partial_{zz} \sim \lambda_{\theta}^{-2}$ in (2.13) with λ_{u} from (2.7) and λ_{θ} from (2.8), we get

$$Nu \sim Pr^{1/3} Re_L^{1/3} \left(C_T Re_T \right)^{1/3} \equiv \frac{Pr^{1/3}}{C_L^{2/3} Re_L^{1/3}} C_T Re_T.$$
(2.15)

Also the bulk contribution to the thermal dissipation is estimated in an identical manner to the corresponding equation in the GL-theory, only with minor changes to reflect the new dependence on C_L , C_T and Re_T . For $\lambda_{\theta} > \lambda_u$

$$\epsilon_{\theta,bulk} \sim \Delta^2 \frac{U_L}{H} \sim \kappa \frac{\Delta^2}{H^2} Pr Re_L \equiv \kappa \frac{\Delta^2}{H^2} \frac{Pr}{C_L} \left(C_T Re_T \right), \qquad (2.16)$$

and for $\lambda_{\theta} < \lambda_{u}$, where the relevant velocity scale is $U_{L}\lambda_{\theta}/\lambda_{u}$, we get

$$\epsilon_{\theta,bulk} \sim \Delta^2 \frac{U_L}{H} \frac{\lambda_\theta}{\lambda_u} \sim \kappa \frac{\Delta^2}{H^2} Pr C_L R e_L^2 N u^{-1} \equiv \kappa \frac{\Delta^2}{H^2} \frac{Pr}{C_L} \frac{(C_T R e_T)^2}{N u}.$$
 (2.17)

We expect (2.14)–(2.17) to be valid in both buoyancy-dominated and shear-dominated regimes. This is accommodated by the fact that the dependencies $C_T(Re_T)$ and $C_L(Re_L)$ behave differently in the buoyancy-dominated and shear-dominated regimes. Compatibility of these relations in the limiting regimes is explained in the subsequent § 2.4.

2.4. Limiting regimes

It is important to note that there exist no pure power laws for Nu, Re_L and C_S as functions of Ra, Pr and Re_S . Nonetheless, it is useful to study the pure scaling power laws that arise from limiting regimes, in the interest of understanding the physics of the system. Based on the dominance of boundary layer and bulk contributions to the kinetic and thermal dissipation rates, the GL-theory provides four regimes I, II, III and IV for the pure RB system. Furthermore, each regime can be divided into two subregimes based on whether the thermal boundary layer is nested into the kinetic boundary layer or *vice versa*. We find that this classification of regimes is also applicable to buoyancy-dominated sheared RB. The phase space of Ra, Pr and Re_S is divided by four different transitions – (i) transition from boundary layer dominated regimes to bulk-dominated regimes, (ii) transition from buoyancy-dominated regime to shear-dominated regime, (iii) transition between regimes where the thermal boundary layer is nested inside the kinetic boundary layer or *vice versa* and (iv) transition from a laminar Prandtl–Blasius-type boundary layer to a turbulent Prandtl–von Kármán-type boundary layer.

It should also be noted that some of these limiting regimes may not exist for sheared RB in the shear-dominated state. For example, at high shear forcing, the shear term in $\epsilon_{u,BL}$ always dominates ϵ_u because shear forcing primarily increases the boundary layer

contribution of the kinetic dissipation $(\nu^3/H^4)C_SRe_S^3$. Since the friction coefficient C_S associated with the applied shear can become independent of Re_S only asymptotically at infinite Re_S , shear-dominated systems with $\epsilon_u \sim \epsilon_{u,bulk}$ cannot exist.

We now proceed to first analyse the shear-dominated regime with $Re_S/Re_L \gg 1$. This can be realised in two ways – namely either Re_L is small or Re_L is not necessarily small but $Re_S \gg Re_L$. When Re_L is small, we assume the existence of a laminar Prandtl–Blasius-type boundary layer with

$$C_L \sim R e_I^{-1/2},$$
 (2.18)

which, along with (2.7) and (2.18), gives

$$C_S \sim Re_L^{1/2} Re_S^{-1}.$$
 (2.19)

With the assumption (2.18), we see that (2.14) associated with small Pr becomes

$$Nu \sim Pr^{1/2}C_T Re_T, \tag{2.20}$$

and (2.15) associated with large *Pr* becomes

$$Nu \sim Pr^{1/3}C_T Re_T. \tag{2.21}$$

For the limiting case of $Re_L = 0$ one can see that relations (2.20) and (2.21) recover the scaling laws for passive transport in Couette (Yerragolam *et al.* 2022*a*) or Poiseuille flow (Kays & Crawford 1993) where the relation between C_T and Re_T depends on whether the kinetic boundary layer is laminar or turbulent. In the presence of a laminar boundary layer, the trivial scaling, $C_T \sim Re_T^{-1}$, applies. When the boundary layer turns turbulent with increased shear forcing, the relation between C_T and Re_T is given by Prandtl (1932) friction law obtained from the log-law mean velocity profile which states

$$\sqrt{\frac{2}{C_T}} = \frac{1}{k} \ln \left(R e_T \sqrt{\frac{C_T}{8}} \right) + B, \qquad (2.22)$$

with $k \approx 0.41$ (Pirozzoli, Bernardini & Orlandi 2014) being the von Kármán (1934) constant and $B \approx 5$ (Pirozzoli *et al.* 2014) indicating the log-law intercept.

When Re_L is not necessarily small and $Re_S \gg Re_L$, we consider that the passive transport relations of (2.20) or (2.21) remain relevant, and that the dependence of $C_T(Re_T)$ is unchanged to that described above. For this case of larger Re_L , we assume that the thermal dissipation is dominated by contributions from the bulk, so that the global dissipation relation (2.1) can be estimated by (2.16) or (2.17). When we compare the passive transport relations with the dissipation estimates, we find that C_L must become independent of Re_L , with $C_L \sim Pr^{1/2}$ for low Pr and $C_L \sim Pr^{1/3}$ for high Pr. Furthermore, the behaviour of Re_L in these cases can be revealed by combining the passive transport relations (2.20) or (2.21) with the boundary-layer estimate for the kinetic dissipation rate (2.10). For low Pr this produces $Re_L \sim Ra^{1/2}Pr^{-3/4}$, and for high Pr we get $Re_L \sim Ra^{1/2}Pr^{-5/6}$, which exactly match the Reynolds number scaling relations found in the boundary layer dominated regime I of the GL-theory for classical RB convection.

Now, we focus on the buoyancy-dominated regimes where $Re_S/Re_L \leq 1$ and $C_SRe_S^3 \ll (Nu-1)RaPr^{-2}$ such that the shear contribution to the kinetic dissipation can be neglected in comparison with the buoyancy contribution. With this restriction, we

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Regime	ϵ_u	$\epsilon_{ heta}$	BL ratio	Nu/Nu_R	<i>KeL/KeR</i>
I_l	BL (2.24)	BL (2.14)	$\lambda_{\theta} > \lambda_{u}$	$(C_L^2 R e_L)^{-1/8}$	$(C_L^2 R e_L)^{-1/4}$
I_u	BL (2.24)	BL (2.15)	$\lambda_{\theta} < \lambda_{u}$	$(C_L^2 Re_L)^{1/12}$	$(C_L^2 R e_L)^{-1/6}$
II_l	bulk (2.11)	BL (2.14)	$\lambda_{\theta} > \lambda_{u}$	$(Re_T/Re_L)^{-1/5}$	$(Re_{T}/Re_{L})^{-2/5}$
II_u	bulk (2.11)	BL (2.15)	$\lambda_{\theta} < \lambda_{u}$	$(C_L^2 Re_L)^{1/5} (Re_T/Re_L)^{-1/5}$	$(C_L^2 R e_L)^{1/5} (R e_T / R e_L)^{-2/5}$
III_l	BL (2.24)	bulk (2.16)	$\lambda_{\theta} > \lambda_{u}$	$(C_L^2 R e_L)^{-1/3}$	$(C_L^2 R e_L)^{-1/3}$
III_u	BL (2.24)	bulk (2.17)	$\lambda_{\theta} < \lambda_{u}$	$(C_L^2 R e_L)^{1/7}$	$(C_L^2 R e_L)^{-1/7}$
IV_l	bulk (2.11)	bulk (2.16)	$\lambda_{\theta} > \lambda_{u}$	$(Re_T/Re_L)^{-1/2}$	$(Re_T/Re_L)^{-1/2}$
IV_u	bulk (2.11)	bulk (2.17)	$\lambda_{\theta} < \lambda_{u}$	$(C_I^2 R e_L)^{1/3} (R e_T / R e_L)^{-1/3}$	$(C_L^2 R e_L)^{1/9} (R e_T / R e_L)^{-4/9}$

Table 2. Scaling relations for the Nusselt number Nu and LSC Reynolds number Re_L in the buoyancy-dominated regime of sheared RB convection with a laminar Prandtl–Blasius-type boundary layer. The first column indicates the GL regime, the second column indicates the bulk or boundary layer (BL) dominance of the kinetic dissipation rate with the applicable scaling estimate in the parenthesis and the third column indicates the bulk or BL dominance of the thermal dissipation rate with the applicable scaling estimate in the parenthesis. The fourth column indicates whether the kinetic BL is thicker than the thermal boundary layer or *vice versa*. The fourth and fifth columns indicate the scaling relations for Nu/Nu_R and Re_L/Re_R using the values of Nu and Re_R estimated for the pure RB system from GL-theory.

approximate (2.2) as

$$\epsilon_u \approx \frac{\nu^3}{H^4} (Nu - 1) Ra P r^{-2}, \qquad (2.23)$$

and (2.10) as

$$\epsilon_{u,BL} \approx \frac{\nu^3}{H^4} C_L R e_L^3. \tag{2.24}$$

Using these approximations, we can provide scaling relations between Nu/Nu_R and Re/Re_R for buoyancy-dominated sheared RB system with Nu_R and Re_R being the Nusselt number and Reynolds number associated with the LSC rolls for the pure RB system. Following the GL-theory, there are various regimes that can be relevant depending on whether the dissipation rates are dominated by boundary layer or bulk contributions, and whether the thermal BL is thicker than the kinetic BL. For each of these regimes, we combine (2.7) and (2.23) with the relevant estimates for the dominant dissipation rate contribution to give expressions for $Nu(Ra, Pr, C_L, Re_L, Re_T)$ and $Re_L(Ra, Pr, C_L, Re_L, Re_T)$. Since our approach is consistent with the GL-theory, the Ra and Pr dependence simply recovers the scaling relations $Nu_R(Ra, Pr)$ and $Re_R(Ra, Pr)$ found for the various regimes of pure RB, providing us with scaling relations for Nu/Nu_R and Re_L/Re_R that only depend on C_L , Re_L and Re_T . In table 2, we outline the relevant to note that these scaling relations are only applicable to buoyancy-dominated sheared RB flows with scaling-wise laminar Prandtl–Blasius-type BLs.

In the buoyancy-dominated classical GL regimes, we can consider the Prandtl–Blasius scaling (2.18) to hold for small shear forcing (i.e. $Re_S \leq Re_L$). With this assumption, $C_L^2 Re_L \approx 1$, so the values of Nu and Re_L remain unchanged for buoyancy-dominated regimes I and III, whereas for buoyancy-dominated regimes II and IV, the non-monotonic behaviour of Nu with increasing Re_S becomes apparent. Although Nu seems to decrease with increasing Re_S in the buoyancy-dominant II and IV regimes, it is important to note that this behaviour is subject to the condition that the BL is a laminar one of

the Prandtl–Blasius type. If the BL becomes turbulent, the expected decrease in Nu in the buoyancy-dominated regime might disappear. In this study, we will explore the Nu response in the buoyancy-dominated II_u regime where we can still observe the decrease in Nu with increasing Re_S within reasonable computational cost.

3. Results from the direct numerical simulations

3.1. Scheme and procedure

In this section, we will compare the scaling relations derived in the previous section against the results from our DNS. Equations (1.1a,b) and (1.2) are solved numerically using the in-house open-source code 'AFiD', which is based on a second-order finite-difference scheme (van der Poel et al. 2015). The code has been extensively validated (Verzicco & Orlandi 1996; Verzicco & Camussi 1997; Stevens, Verzicco & Lohse 2010; Stevens, Lohse & Verzicco 2011; Kooij et al. 2018). We impose periodic boundary conditions in the horizontal directions and no-slip boundary conditions at the top and bottom walls. For most simulations, we use domains of aspect ratios $\Gamma_x = 8$ and $\Gamma_y = 4$. We also performed CRB simulations with $\Gamma_x = 48$ and $\Gamma_y = 24$ for $Ra = 10^7$, Pr = 1 to study large-scale flow structures. Due to the need for high resolution at large Ra, the RB simulations for $Ra = 10^{10}$, Pr = 1 and $Ra = 10^{11}$, Pr = 1 were performed in domains of aspect ratios $\Gamma_x = \Gamma_y = 4$, while the RB simulation for $Ra = 10^{12}$, Pr = 1 was performed in domain of aspect ratios $\Gamma_x = \Gamma_y = 2$. For the CRB simulations, the wall velocities $(-U_w)$ and U_w were imposed as Dirichlet boundary conditions on the bottom and top walls, respectively. This is done for numerical reasons (Bernardini et al. 2013) and does not affect the analysis of the results. The equivalent velocity fields of the CRB system with the bottom wall at rest and the top wall at $2U_w$ can be obtained by a simple Galilean transformation, i.e. by adding U_w to the numerically obtained flow field. For the PRB simulations, the volume forcing term Π is computed at each time step to ensure a constant mass flow rate (Quadrio et al. 2016).

We use a uniform discretisation in the horizontal, periodic directions and a non-uniform grid in the wall-normal direction, in which we employ higher grid resolution in the BLs next to the walls. The thermal BL was ensured to be sufficiently resolved according to the resolution requirements put forward by Shishkina *et al.* (2010). The near-wall resolution is comparable to that of Lozano-Durán & Jiménez (2014); Pirozzoli *et al.* (2014); Lee & Moser (2018) to ensure that the kinetic BL is sufficiently resolved. The simulations were run for a long enough physical time for the standard deviation of *Nu* to converge within approximately 1% of its mean value. In our previous work (Yerragolam *et al.* 2022*b*), we verified that the *Nu* and Re_{τ} obtained from a domain with $\Gamma_x = 8$ and $\Gamma_y = 4$ shows a difference of less than 1% from the *Nu* obtained from a domain with $\Gamma_x = 48$ and $\Gamma_y = 24$. This observation is also supported by the fact that *Nu* for the RB system converges at an approximate aspect ratio $\Gamma_x = \Gamma_y = 4$ (Stevens *et al.* 2018).

Since many of the scaling relations in § 2.4 rely on the value of the wind Reynolds number Re_R , obtaining an estimate for Re_R from the numerical simulations of pure RB is necessary. The value of Re_R can be estimated in two possible ways. The first estimate can be obtained by using $Re_{max} = U_{max}H/v$ where U_{max} is the maximum value of the root mean squared (r.m.s.) horizontal velocity profile $u_R(z) \equiv \sqrt{\langle u_x^2 + u_y^2 \rangle_{A,t}}$ at a height $z = \delta$ as shown in figure 2(a). The second estimate can be obtained by using the global r.m.s. velocity $Re_{rms} = U_{rms}H/v$ with $U_{rms} \equiv \sqrt{\langle u_x^2 + u_y^2 + u_z^2 \rangle_{V,t}}$. In figure 2(b) and 2(c), we can see that these estimates are almost identical, providing strong evidence that the LSC



Figure 2. (a) The r.m.s. horizontal velocity profile $u_R(z)$ normalised with the free-fall velocity for RB system with $Ra = 10^7$, Pr = 1.0 is shown as a function of wall-normal height z, indicating the maximum value U_{max} occurring at a wall-normal height δ . (b) Value of Re_{max} obtained from U_{max} is plotted against the globally averaged r.m.s. velocity Re_R of the RB system. The dashed grey line indicates $Re_{max} = Re_{rms}$, showing that these two estimates are virtually identical. (c) Re_{max} compensated with Re_{rms} plotted against Re_{rms} , once again emphasising that the two estimates are equivalent within the 5 decades of Ra and nearly 3 decades of Re_{rms} investigated here. The values of Ra are shown on the top for reference. This figure is also available at https:// www.cambridge.org/S0022112024008723/JFM-Notebooks/files/figure2.

rolls driving the wind at the wall also provide the dominant contribution to the mean kinetic energy in RB convection. In all the results discussed henceforth, we adopt Re_{rms} as an estimate for Re_R . Unlike in Couette or Poiseuille flow, the mean shear stress at the wall is zero in RB convection. However, we can use the mean gradient of the r.m.s. horizontal velocity $\langle \partial_z u_R(z) \rangle_{W,t}$ to calculate the friction coefficient C_R associated with the large-scale circulation. Here, $\langle \cdots \rangle_{W,t}$ indicates time averaging over the surface of the walls. The variation of C_R with Re_R is discussed separately in § 3.5.

3.2. Global response parameters

We now validate the scaling relations for Nu and C_S derived in § 2.4. Within the parameter range of Ra, Pr and Re simulated, we already observe multiple transitions. As shear forcing is increased, we undergo transition from the buoyancy-dominated regime to the shear-dominated regime. For low shear forcing, $\epsilon_u \sim \epsilon_{u,bulk}$ with $\lambda_{\theta} < \lambda_u$ whereas for high shear forcing, $\epsilon_u \sim \epsilon_{u,BL}$ with $\lambda_{\theta}/\lambda_u \sim Pr^{1/2}$ (Yerragolam *et al.* 2022*a*). Additionally, at low shear forcing, we observe a laminar Prandtl–Blasius-type BL which undergoes a transition into a turbulent one at high shear.

In the buoyancy-dominated regime, we assume the presence of Prandtl–Blasius-type kinetic boundary layer with the friction coefficient C_L given by (2.18). For the parameter range of our simulations, the relevant convection regime is II_u , so combining the relevant relation from table 2 with (2.18), we arrive at

$$\frac{Nu}{Nu_R} \sim \left(\frac{Re_T}{Re_L}\right)^{-1/5}.$$
(3.1)

In the buoyancy-dominated regimes, we also take (2.19) for the friction coefficient C_S associated with the imposed shear. In order to further simplify these equations, we

approximate $Re_L \approx Re_R$ in the buoyancy-dominated regime, which is justified by the weak variation of Re_L with increasing Re_S in (table 2). With this assumption, we can rewrite (2.19) as

$$C_S \sim \sqrt{Re_R}/Re_S,\tag{3.2}$$

and use (2.5) to rewrite (3.1) as

$$Nu/Nu_R \sim \left(\sqrt{1 + \left(Re_S^2/Re_R^2\right)}\right)^{-1/5}.$$
(3.3)

These equations show good agreement with the numerical data plotted in figures 3(a)-3(d) for the buoyancy-dominated regime. Note that (3.1) and (3.3) do not explicitly state the dependence of $Nu_R(Ra, Pr)$ or $Re_R(Ra, Pr)$. There is no pure scaling exponent for $Nu_R(Ra)$ in pure RB for regime II_u . The values of Nu_R and Re_R used for figures 3(b) and 3(d) are obtained from numerical simulations of pure RB. The present extension to the GL-theory assumes that the values of Nu_R and Re_R are known a priori, and only attempts to provide scaling relations for the normalised quantities $Nu/Nu_R(Re_S/Re_R)$ and $C_S/C_R(Re_S/Re_R)$ in the buoyancy-dominated regime.

For the shear-dominated regime, we observe that the BL becomes turbulent. In this case, (2.19) is no longer valid. Instead, the relation between C_T and Re_T is given by (2.22). Note that, in the limiting case of very high-shear forcing, $Re_T \approx Re_S$. Equation (2.22) can be rewritten as

$$\sqrt{\frac{2}{C_S}} = \frac{1}{k} \ln\left(Re_S \sqrt{\frac{C_S}{8}}\right) + B,\tag{3.4}$$

and (2.20) can be rewritten using (2.7) as

$$Nu \sim Pr^{1/2}C_S Re_S, \tag{3.5}$$

which agrees well with the numerical data points in figures 3(c) and 3(d) at very high shear forcing. However, it is more useful to substitute the value of C_T obtained from (2.22) into (2.20) and approximate Re_T from (2.5) as

$$Re_T \approx \sqrt{Re_R^2 + Re_S^2},\tag{3.6}$$

to obtain the dashed lines plotted in figure 3(a), which show better agreement for a larger range of Re_S/Re_R in the shear-dominated regime. The approximation given by (3.6) is then validated in figure 4(a). The additional energy term $\gamma C_S Re_S^2$ from (2.4) that corresponds to the turbulent fluctuations arising from shear forcing is shown in figure 4(b) with the value of the prefactor $\gamma \approx 10.24$. Note that the contribution from the fluctuations is much smaller than the contribution from the mean streamwise velocity, thereby making (3.6) a good approximation.

3.3. Large-scale convection rolls

Next, we confirm the theoretical assumptions on the LSC rolls made in § 2.2. The three-dimensional volume visualisation in figure 5 shows the streamlines associated with these LSC rolls and it can be seen that they are predominantly oriented in the spanwise direction. However, the large-scale temperature structures in figure 5 are seen to be aligned neither fully along the streamwise direction, nor fully along the spanwise direction but along a diagonal. The thermal plumes that comprise these large-scale structures experience





Figure 3. (a) Value of $Nu/Pr^{1/2}$ plotted against Re_S . The black solid line indicates (3.5), the dashed lines indicate (2.20) and the dotted lines indicate (3.3). (b) Value of Nu/Nu_R plotted against Re_S/Re_R . Dotted lines indicate (3.3). (c) Value of C_S plotted against Re_S . The black solid line indicates (3.4) and coloured dashed lines indicate (2.19). (d) Value of C_S normalised with C_R plotted against Re_S/Re_R . Coloured dashed lines indicate (3.4) and black dashed line indicates (2.19). The data for PRB are indicated with plus markers and the data for CRB are indicated with dot markers. This figure is also available at https://www.cambridge.org/S0022112024008723/JFM-Notebooks/files/figure3.

the advective effects of both U_L in the spanwise direction and U_S in the streamwise direction. Therefore, the orientation of these large scale temperature flow structures in the x-y plane is tilted along a diagonal whose slope is approximately given by U_L/U_S . This is made clearer in figure 6 through the visualisation of the non-dimensional time-averaged local shear stress $\tau'_w \equiv (\tau'_x, \tau'_y, 0)$ computed at the bottom wall in the large aspect ratio CRB system by subtracting the wall-averaged streamwise shear stress in the following way:

$$\mathbf{t}_{x}' = HU_{F}^{-1} \left\langle \partial_{z} u_{x} - \left\langle \partial_{z} u_{x} \right\rangle_{x,y,t} \right\rangle_{t}, \quad \mathbf{\tau}_{y}' = HU_{F}^{-1} \left\langle \partial_{z} u_{y} \right\rangle_{t}. \tag{3.7a,b}$$



Figure 4. (a) Value of Re_T/Re_R plotted against Re_S/Re_R . The black solid line indicates the relation (2.5). (b) Value of $(\sqrt{Re_T^2 - Re_S^2})/Re_R$ plotted against $\sqrt{C_S}Re_S/Re_R$. The black dashed line indicates the relation (2.4) with $\gamma \approx 10.24$. The data for PRB are indicated with plus markers and the data for CRB are indicated with dot markers. This figure is also available at https://www.cambridge.org/S0022112024008723/JFM-Notebooks/files/figure4.

Figure 6(*a*) shows the wall shear for the RB system i.e. for $Re_w = 0$ and the zoomed inset in figure 6(*b*) shows the vectors of τ'_w . As expected, the LSC rolls are randomly oriented and no global alignment of τ'_w is observed. A visual inspection of figures 6(*c*), 6(*e*) and 6(*g*) reveals that the large-scale flow structures seem to be oriented along the diagonal whose slope is approximately given by Re_R/Re_w . However, figures 6(*d*), 6(*f*) and 6(*h*) show that τ'_w is primarily oriented along the spanwise direction in the transitional regime. Figure 6(*i*) shows the breakdown of the LSC rolls and the formation of large meandering flow structures (Hutchins & Marusic 2007; Blass *et al.* 2020, 2021) in the shear-dominated regime, while figure 6(*j*) shows that τ'_w is predominantly aligned in the streamwise direction in the shear-dominated regime.

For further confirmation of the changes in the LSC rolls, we study the probability distribution function $\phi(\alpha)$ of the angle α spanned by the horizontal velocity component fluctuations $u'_{h} \equiv (u'_{x}, u'_{y}, 0)$ with the streamwise direction x. In figures 7(a)-7(e) it can be seen that the behaviour of $\phi(\alpha)$ is qualitatively quite similar for the CRB and PRB systems. In the RB system, the LSC rolls are randomly oriented as shown by a uniform $\phi(\alpha)$ in figure 7(a). In the buoyancy-dominated regime, the LSC roll has a strong tendency to align in the spanwise directions in the transitional regime, as shown in figures 7(b) and 7(d). In the shear-dominated regime, the velocity fluctuations are predominantly aligned in the streamwise direction as in the case of turbulent Couette/Poiseuille flows, as seen in figures 7(c) and 7(e). At the thermal BL height, the symmetry of $\phi(\alpha)$ about $\alpha = \pm \pi/2$ is strongly suggested for all three regimes by the data from the numerical simulations as shown in figures 7(f)-7(h), confirming the assumption made in (2.6). In figure 7(f), the small non-uniformity in $\phi(\alpha)$ is attributed to the numerical confinement experienced by the flow structures in domains of a smaller aspect ratio of $\Gamma_x = 8$ and $\Gamma_y = 4$. For the relatively unconfined case with $\Gamma_x = 48$ and $\Gamma_y = 24$, the probability distribution is nearly uniform for all values of α .



Figure 5. Three-dimensional volume visualisation of the time-averaged reduced temperature field of a CRB system with $Ra = 10^7$, Pr = 1.0, $Re_w = 1414$, $\Gamma_x = 48$, $\Gamma_y = 24$ with the averaging time being 100 free-fall time units. The green curves indicate the streamlines of the time-averaged velocity field $\langle u \rangle_l - U_w \hat{x}$. The spanwise reorientation and streamwise sweeping of the plumes are evident. The zoomed inset shows the time-averaged two-dimensional visualisation of the temperature and velocity vectors in the *x*–*z* plane. The plumes carry streamwise momentum along with temperature and are swept in the streamwise direction, causing a reduction in the heat transport.

3.4. Dissipation rates

We now investigate the bulk and BL contributions to the global kinetic and thermal dissipation rates described in § 2.1. We validate the rigorous relations given by (2.1) and (2.2) using the data obtained from numerical simulations as shown in figures 8(*a*) and 8(*b*), respectively. For the range of *Ra* and *Pr* studied in this work, as long as the flow is buoyancy dominated, the system is in the I_u regime with $\epsilon_u \sim \epsilon_{u,bulk}$. With increasing shear forcing, the shear term of (2.2) increases the BL contribution of the kinetic dissipation due to the formation of streamwise velocity gradients close to the wall. For sufficiently strong shear, the kinetic dissipation will be dominated by the BL contribution rate is dominated by the BL contribution for the entire range of Re_S considered in this study but the contribution reduces noticeably towards higher Re_S . For extremely strong shear, a possibility of a transition towards bulk dominance in thermal dissipation cannot be ruled out.

3.5. Friction coefficient in Rayleigh-Bénard flow

So far, we have considered only the laminar Prandtl-Blasius-type kinetic BLs in the buoyancy-dominated regime with the BL only becoming turbulent in highly



Figure 6. Visualisation of τ'_{w} for the CRB system with $Ra = 10^{7}$ and Pr = 1. Panels (c), (e) and (g) are in the buoyancy-dominated regime and panel (i) is in the shear-dominated regime. The colour bars indicate the magnitude of τ'_{w} while the black arrows in the magnified panels (b), (d), (f), (h) and (j) indicate the direction of τ'_{w} . The black lines in (c-h) indicate the slope of $(Re_{w}/Re_{R})^{-1}$. This figure is also available at https://www. cambridge.org/S0022112024008723/JFM-Notebooks/files/figure6.



Figure 7. The probability distribution $\phi(\alpha)$ of the flow orientation angle α for all heights $0 \le z/\lambda_{\theta} \le Nu$ for $Ra = 10^7$, Pr = 1, $\Gamma_x = 48$ and $\Gamma_y = 24$ in (a) RB flow (i.e. with $Re_S = 0$), (b,d) buoyancy-dominated regime with $Re_S/Re_R \approx 2.11$ and (c,e) shear-dominated regime with $Re_S/Re_R \ge 10$. Panels (b,c) are for the CRB system while panels (d,e) are for the PRB system. The black dashed lines indicate the height of the thermal BL and the colour bar indicates the magnitude of the probability. The probability distribution $\phi(\alpha)$ at the thermal BL height plotted against angle α for various Ra and Pr in (f) RB flow, (g) buoyancy-dominated regime and (h) shear-dominated regime. The solid lines are the CRB system with $\Gamma_x = 48$ and $\Gamma_y = 24$, the dashed lines are for the CRB system with $\Gamma_x = 8$ and $\Gamma_y = 4$. This figure is also available at https://www.cambridge.org/S0022112024008723/JFM-Notebooks/files/figure7.

shear-dominated regime. However, we now consider the possibility of turbulent kinetic BLs even in pure RB flow which corresponds to the so-called 'ultimate' regime in RB flow (Kraichnan 1962; Grossmann & Lohse 2011; Roche 2020; Lohse & Shishkina 2023). While the existence of logarithmic temperature profiles has already been observed at high *Ra* (Grossmann & Lohse 2012; Ahlers *et al.* 2012; Ahlers, Bodenschatz & He 2014), it is yet to be seen if logarithmic behaviour given by

$$u_R^+(z^+) = \frac{1}{k}\log(z^+) + B,$$
(3.8)



Figure 8. (a) Ratio $(H^2\epsilon_{\theta}\kappa^{-1}\Delta^2)/Nu$ showing good agreement with (2.1) indicated using the solid black line, and (b) ratio $(H^4\epsilon_u/v^3)/((Nu-1)RaPr^{-2} + C_SRe_3^S)$ plotted against Re_S showing good agreement with (2.2) of the manuscript shown using the solid black line. (c) Boundary layer contribution to the global kinetic dissipation rate $\epsilon_{u,BL}$ and (d) the BL contribution to the global thermal dissipation rate $\epsilon_{\theta,BL}$. The circle markers are for the CRB system, while plus markers are for the PRB system. This figure is also available at https://www. cambridge.org/S0022112024008723/JFM-Notebooks/files/figure8.

with

$$u_{R}^{+}(z^{+}) = u_{R}(z)/u_{\tau}, \quad z^{+} = \frac{zu_{\tau}}{2\nu}, \quad u_{R}(z) \equiv \sqrt{\left\langle u_{x}^{2} + u_{y}^{2} \right\rangle_{A,t}}, \quad u_{\tau} = \sqrt{\nu \left\langle \partial_{z} u_{R} \right\rangle_{W,t}},$$
(3.9*a*-*d*)

can be observed in the velocity profiles.

At the highest thermal forcing studied in this work with $Ra = 10^{12}$, we start to also observe some hints of what could possibly be the onset of a log layer (see figure 9*a*), although it cannot be conclusively confirmed with the currently available data. Assuming that such a log layer could exist, an estimate of the modified von Kármán constant *k* is obtained from the inflection point of the diagnostic function plotted in figure 9(*b*), giving



Figure 9. (a) Value of u_R^+ plotted against z^+ , showing the onset of a log layer with the grey dashed line. The velocity profile for $Ra = 10^{12}$, Pr = 1.0 is indicated in magenta. (b) Diagnostic function plotted against z^+ with the inflection point and the corresponding von Kármán constant indicated. (c) Value of δ^+ plotted against Re_R shows good agreement with $\delta^+ = (1/2)Re_R^{1/2}$, indicated with the dashed grey line. (d) Value of C_R plotted against Re_δ showing a collapse without any prefactors. The solid grey line indicates $C_R \sim 4Re_{\delta}^{-2/3}$ and the dashed grey line indicates the modified Prandtl (1932) friction law given by (3.10*a*,*b*). (e) Value of C_R/a plotted against Re_R , showing the Prandtl–Blasius scaling (2.18) at low Re_R with the solid grey line and modified Prandtl (1932) friction law given by (3.11*b*) plotted against Re_R . The solid grey line indicates $C_{\epsilon} \sim Re_R^{-1/2}$ for the regime with $\epsilon_u \sim \epsilon_{u,BL}$, the dashed grey line indicates the modified Prandtl (1932) friction law (3.10*a*,*b*). At higher Re_R , $\epsilon_u \sim \epsilon_{u,bulk}$, leading to C_{ϵ} becoming independent of Re_R , as indicated with the dash-dotted and dotted grey lines. This figure is also available at https://www.cambridge.org/S0022112024008723/JFM-Notebooks/files/figure9.

 $k \approx 1$. Correspondingly, the intercept $B \approx 5.2$ is found by fitting the data, as shown in figure 9(*a*).

In the presence of such a logarithmic layer, we can now hypothesise about a relation for C_R which is analogous to relation (3.4) for C_S and (2.22) for C_T by stating that in the limit of highly turbulent BL, $u_R^+ = U_R$ at $z^+ = \delta^+$, where δ is the wall-normal distance to the peak velocity shown in figure 4. This gives us

$$\sqrt{\frac{2}{C_R}} = \frac{1}{k} \ln \left(Re_{\delta} \sqrt{\frac{C_R}{8}} \right) + B, \quad Re_{\delta} = U_R \delta / \nu.$$
(3.10*a*,*b*)

If applicable to high *Ra* RB flow, the general form of this equation could suggest universality in the behaviour of wall-bounded flows even though the values of *k* and *B* might be different from those observed for pipe or channel flows. By plotting δ^+ against

 Re_R , we find a very good fit with the scaling $\delta^+ \sim Re_R^{1/2}$, as shown in figure 9(c). At present, we can only provide this scaling empirically because more data at extremely high Ra are needed to understand the dynamics of the turbulent BL which is difficult due to the high computational expense of such numerical simulations. Plotting C_R against Re_δ in 9(d), we find a nice collapse of the data for different aspect ratios and geometries. At low Re_R , C_R seems to scale as $Re_{\delta}^{-2/3}$, which can be obtained from the empirical scaling $\delta^+ \sim Re_R^{1/2}$ shown in figure 9(c) and the assumption of a Prandtl–Blasius-type scaling of $C_R \sim Re_R^{-1/2}$. At high Re_R , we see good qualitative agreement between the data and the modified Prandtl (1932) friction law given by (3.10*a*,*b*).

We can further investigate the behaviour of C_R by computing

$$C_R \equiv a R e_R^{-1/2}, \tag{3.11a}$$

$$C_{\epsilon} \equiv b(Nu-1)RaPr^{-2}Re_R^{-3}, \quad \begin{cases} C_{\epsilon} \sim C_R & \text{if } \epsilon_u \sim \epsilon_{u,BL} \\ C_{\epsilon} = \text{const.} & \text{if } \epsilon_u \sim \epsilon_{u,bulk}, \end{cases}$$
(3.11b)

where *a* and *b* are prefactors that are obtained by fitting the data from figures 9(*e*) and 9(*f*), accounting for the effects of aspect ratio and geometry. Note that, when the kinetic dissipation of the RB system is dominated by the contribution from the BL, $C_{\epsilon} \sim C_R$ in (3.11*b*). The data points at lower Re_R in figure 9(*e*) are observed to follow (2.18), which is consistent with the Prandtl–Blasius scaling. For higher Re_R , corresponding to $Ra > 10^{10}$, C_R shows better agreement with the Prandtl (1932) friction law given by (3.10*a*,*b*). Figure 9(*f*) shows a very similar behaviour to figure 9(*e*) but, in addition, we observe that C_{ϵ} becomes independent of Re_R at very high values of Re_R . Although this does not reflect the true dependence of C_R on Re_R , this apparent dependence is expected because the kinetic dissipation of the RB system undergoes a transition from being dominated by the BL to being dominated by the bulk (Lohse 1994). It should also be noted that this transition occurs at higher Re_R for the more confined cylindrical RB simulations because more kinetic driving is required to overcome the viscous dissipation in the additional BLs on the side walls that are not present in the periodic box RB simulations (Ahlers *et al.* 2022).

4. Conclusions

In summary, we have developed a framework by extending the GL-theory for RB turbulence to sheared RB turbulence. As in the case of RB flow, we observe that there are no pure scaling exponents for the Nusselt number Nu and the friction coefficient C_S . This also holds for high thermal or shear driving where the BLs no longer obey scaling relations associated with the Prandtl–Blasius (Prandtl 1904; Blasius 1908) BL theory but start to become more turbulent. In such cases, we observe that the relation for $C_S(Re_S)$ is well described by the friction law of Prandtl (1932). In addition, we find that a modified version of the Prandtl's (1932) friction law for the large-scale circulation rolls $C_L(Re_L)$ analogous to $C_S(Re_S)$ agrees well with the DNS data, suggesting some sort of universality in the relation between the shear stress and the flow velocity that generates that shear.

It is also interesting to note that the relations are identical for CRB or PRB systems once the appropriate velocity scale is chosen as a control parameter. This suggests that the flow physics is not strongly affected by the geometry of the system or by the way in which shear forcing is applied. The flow characteristics of these systems are essentially determined by the ratio of shear driving to thermal driving, given by Re_S/Re_R . Applying

shear to the RB system causes increased coherence in the streamwise direction and leads to a re-orientation of the LSC rolls and causes them to align more in the spanwise direction, with the thermal plumes also transporting the momentum imparted by the shear forcing. In the buoyancy-dominated regime with $Re_S \simeq Re_R$, this may lead to enhanced streamwise mixing between hot and cold plumes at a time scale that is smaller than the time scale of heat diffusion at the wall. This leads to heat entrapment in the bulk and a reduction in Nu.

Taking into account the orientation of the LSC rolls and the bulk dominance of ϵ_u in the buoyancy-dominated regime, we show that the orientation of large-scale flow structures can also be predicted to a reasonable degree by the ratio Re_S/Re_R , and we provide scaling relations for the $Nu(Re_S/Re_R)$ and $C_S(Re_S/Re_R)$, which are shown to agree well with the numerical simulations for $10^6 \le Ra \le 10^8$, $0.5 \le Pr \le 5.0$ and $0 \le Re_S \le 10^4$. However, the evidence from the DNS is limited at the moment due to its high computational costs, thereby restricting the parameter range in which the proposed scaling laws can be validated. Simulations for very high or very low values of Pr as well as for high Ra or Re_S can be very demanding, and it remains to be seen if the assumptions made in this work and the extended theory hold well in other control parameter ranges.

Supplementary material. Computational Notebook files are available as supplementary material at https://doi.org/10.1017/jfm.2024.872 and online at https://www.cambridge.org/S0022112024008723/JFM-Notebooks.

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Declaration of interests. The authors report no conflict of interest.

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Appendix. Simulation parameters

In table 3 we provide the physical and numerical input parameters used for the new sheared RB simulations conducted for this study. In addition to the new simulations, data for large aspect ratio simulations with $\Gamma_x = 48$ and $\Gamma_y = 24$ are taken from Yerragolam *et al.* (2022*b*), and data for high *Ra* cylindrical simulations are taken from Hartmann *et al.* (2023).

C_S												1.56	1.09	0.459	0.511	0.399	0.327	0.3	0.251	0.224	0.203	0.102	0.0676	0.0546	0.0442	0.038	0.0332	0.0296	0.0268	0.0261	0.0255	
Nu_{ϵ_u}	4.38	8.32	15.32	15.73	16.02	16.19	30.71	61.98	129.80	287.26	732.83	8.27	8.29	8.25	8.26	8.25	8.29	8.24	8.22	8.18	8.19	7.85	7.47	6.95	6.66	6.45	6.35	6.33	6.26	6.26	6.27	
$N u_{\epsilon_{ heta}}$	4.38	8.31	15.33	14.47	16.04	16.14	30.68	61.86	129.80	288.02	709.50	8.27	8.29	8.26	8.27	8.26	8.29	8.24	8.21	8.19	8.18	7.84	7.47	6.97	6.67	6.46	6.35	6.31	6.28	6.25	6.27	
Nu_{w}	4.38	8.31	15.32	15.75	16.05	16.14	30.68	61.83	129.78	287.97	709.21	8.26	8.29	8.26	8.28	8.26	8.29	8.25	8.21	8.19	8.18	7.84	7.47	6.97	6.66	6.46	6.35	6.31	6.27	6.25	6.27	
Re_S	0	0	0	0	0	0	0	0	0	0	0	10	20	30	40	50	60	70	80	90	100	200	300	400	500	600	700	800	900	920	940	
Pr	1.0	1.0	0.5	1.0	3.0	5.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	
Ra	10^{5}	10^{6}	10^{7}	10^{7}	10^{7}	10^{7}	10^{8}	10^9	10^{10}	10^{11}	10^{12}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	page.
N_{BL}	41	40	29	28	42	42	29	28	28	24	23	40	40	40	40	40	40	40	40	40	40	41	42	43	44	45	46	46	46	46	46	ee next j
z^{+}_{z}												0.05	0.09	0.09	0.12	0.14	0.15	0.17	0.17	0.18	0.19	0.28	0.34	0.40	0.45	0.51	0.55	0.60	0.64	0.64	0.65	caption s
+* **												0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.01	0.01	0.01	3. For (
y_+												0.05	0.08	0.07	0.11	0.12	0.13	0.14	0.15	0.16	0.17	0.23	0.29	0.34	0.39	0.43	0.47	0.51	0.54	0.55	0.55	Table
^+x												0.05	0.08	0.07	0.11	0.12	0.13	0.14	0.15	0.16	0.17	0.23	0.29	0.34	0.39	0.43	0.47	0.51	0.54	0.55	0.55	
N_z	192	256	256	256	384	384	384	512	768	1024	1536	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	
N_y	512	768	768	768	1024	1024	1024	1536	2048	3072	2048	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	
N_x	1024	1536	1536	1536	2048	2048	2048	3072	2048	3072	2048	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	
Γ_y	4	4	4	4	4	4	4	4	4	4	7	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	
Γ_x	8	8	8	8	8	8	8	8	4	4	7	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	
System	RB	RB	RB	RB	RB	CRB																										

Transport in sheared Rayleigh-Bénard convection

C_S	0.0252	0.0245	0.0241	0.0203	0.0164	0.0125	0.0105	0.00988	0.00957	0.00897	0.00867	0.0079	0.00738	0.00702	0.00652	0.00639	0.00613	0.00602	0.543	3.05	1.14	0.829	0.843	0.554	0.746	0.558	0.398	0.477	0.227	0.16	0.112	
Nu_{ϵ_u}	6.22	6.21	6.21	6.18	6.19	6.25	6.30	6.38	6.73	6.67	6.86	7.82	9.12	10.45	11.51	12.86	13.91	15.17	15.30	15.29	15.33	15.26	15.26	15.22	15.33	15.35	15.24	15.36	15.31	15.23	15.21	
$Nu_{\epsilon_{\theta}}$	6.24	6.19	6.21	6.19	6.19	6.26	6.29	6.38	6.72	6.73	6.92	7.83	9.16	10.46	11.50	12.86	13.92	15.17	15.30	15.29	15.32	15.26	15.27	15.22	15.31	15.35	15.26	15.36	15.31	15.23	15.21	
Nu_w	6.24	6.20	6.21	6.19	6.19	6.26	6.29	6.39	6.72	6.74	6.92	7.83	9.16	10.45	11.51	12.87	13.92	15.18	15.30	15.29	15.32	15.26	15.26	15.22	15.31	15.34	15.26	15.35	15.31	15.24	15.21	
Re_S	096	980	1000	1200	1500	2000	2400	2600	2800	3000	3200	4000	5000	6000	7000	8000	0006	$10\ 000$	10	20	30	40	50	60	70	80	90	100	200	300	400	
Pr	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	
Ra	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^7	10^7	10^7	10^7	10^7	10^7	10^{7}	10^{7}	10^{7}	10^7	10^7	10^{7}	10^{7}	page.
N_{BL}	46	46	46	46	46	46	46	45	44	44	44	41	38	35	33	31	30	29	29	29	29	29	29	29	29	29	29	29	29	29	29	ee next j
$\zeta_m^{+\chi}$	0.66	0.66	0.67	0.74	0.83	0.97	1.06	1.12	1.18	1.23	1.29	1.54	1.86	2.17	2.44	2.76	3.04	3.35	0.03	0.15	0.14	0.16	0.20	0.19	0.26	0.26	0.25	0.30	0.41	0.52	0.58	caption s
z_{w}^{2+}	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.02	0.02	0.02	0.02	0.03	0.03	0.03	0.04	0.04	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	3. For
\mathbf{y}^+	0.56	0.57	0.57	0.63	0.71	0.82	0.91	0.95	1.01	1.05	1.10	1.31	1.58	1.85	2.08	2.35	2.59	2.86	0.03	0.13	0.12	0.13	0.17	0.16	0.22	0.22	0.21	0.25	0.35	0.44	0.49	Table
^+x	0.56	0.57	0.57	0.63	0.71	0.82	0.91	0.95	1.01	1.05	1.10	1.31	1.58	1.85	2.08	2.35	2.59	2.86	0.03	0.13	0.12	0.13	0.17	0.16	0.22	0.22	0.21	0.25	0.35	0.44	0.49	
N^2	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	
N_y	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	
N_x	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	
Γ_y	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	
Γ_x	8	8	8	8	8	×	×	8	×	8	8	×	×	×	×	8	×	×	8	8	×	8	8	8	8	8	×	8	×	8	×	
System	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB																		

1000 A74-24

C_S	.0918 0786	0688	061	055	049	041	035	0272	0185	0141	0117	00994	00873	00787	00729	00689	42	56	76	853	LL	679	391	425	391	338	168	112	0878	0745	
	00	.0	.0	.0	.0	.0	0.	0	.0	0.	.0	.0	.0	0	0	0		.0		.0	0.	.0	0	0.	0	0.	0.	.0	.0	.0	
Nu_{ϵ_u}	15.09 15.08	14.85	14.74	14.57	14.47	14.20	13.63	12.92	12.04	11.65	11.43	11.47	11.67	11.91	12.32	12.82	15.73	15.78	15.78	15.69	15.66	15.69	15.82	15.70	15.74	15.76	15.66	15.60	15.47	15.08	
$N u_{\epsilon_{\theta}}$	15.11 15.06	14.90	14.75	14.58	14.47	14.22	13.65	12.92	12.05	11.64	11.44	11.48	11.67	11.92	12.32	12.83	15.73	15.77	15.77	15.66	15.67	15.69	15.79	15.74	15.76	15.76	15.65	15.57	15.44	15.07	
$ u_w $	5.10 5.06	4.89	4.74	4.58	4.48	4.22	3.64	2.92	2.04	1.64	1.43	1.48	1.66	1.92	2.32	2.82	5.73	5.77	5.77	5.66	5.67	5.68	5.79	5.74	5.76	5.76	5.64	5.58	5.45	5.08	
Z		-		-	-	-	-	-	-	-		-	-		-	-	-	-	-	-	1	-	-	-	-	-	-	-	-	Т	
Re_S	500 600	700	800	900	1000	1200	1500	2000	3000	4000	5000	6000	7000	8000	9000	10000	10	20	30	40	50	60	70	80	90	100	200	300	400	500	
Pr	0.5 0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	
Ra	10^{7}	10^{7}	10^{7}	10^7	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^7	10^7	10^7	10^{7}	10^{7}	10^{7}	10^7	10^7	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	age.
N_{BL}	29 29	29	29	29	30	30	31	31	33	33	33	33	33	33	32	32	28	28	28	28	28	28	28	28	28	28	28	28	29	29	ee next p
z_m^+	0.65 0.73	0.79	0.85	0.91	0.96	1.05	1.21	1.43	1.77	2.06	2.34	2.58	2.83	3.07	3.32	3.59	0.05	0.06	0.18	0.16	0.19	0.21	0.19	0.23	0.24	0.25	0.35	0.43	0.51	0.59	caption se
z_w^{+2}	0.01 0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.02	0.02	0.03	0.03	0.03	0.03	0.04	0.04	0.04	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	3. For 6
y^+	0.56 0.62	0.68	0.73	0.78	0.82	0.89	1.03	1.22	1.50	1.75	1.99	2.20	2.41	2.61	2.83	3.06	0.04	0.06	0.16	0.14	0.16	0.18	0.16	0.19	0.21	0.21	0.30	0.37	0.44	0.50	Table
^+x	0.56 0.62	0.68	0.73	0.78	0.82	0.89	1.03	1.22	1.50	1.75	1.99	2.20	2.41	2.61	2.83	3.06	0.04	0.06	0.16	0.14	0.16	0.18	0.16	0.19	0.21	0.21	0.30	0.37	0.44	0.50	
N_z	256 256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	
N_y	768 768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	
	9 9	90	99	36	99	20	20	99	20	20	36	36	36	99	36	99	90	20	99	99	99	90	99	20	36	20	36	90	90	99	
Ň	153 153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	153	
Γ_y	44	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	
Γ_x	∞ ∞	8	8	8	8	×	×	8	8	8	8	8	8	8	8	8	8	8	8	8	8	×	8	8	8	8	8	8	8	8	
System	CRB CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	

Transport in sheared Rayleigh-Bénard convection

C_S	0.0615	0.0552	0.0497	0.0443	0.0413	0.0306	0.0201	0.0143	0.0105	0.00777	0.00701	1.98	1.03	0.717	0.503	0.469	0.367	0.345	0.296	0.269	0.237	0.12	0.0807	0.0622	0.0555	0.0482	0.0421	0.0372	0.0333	0.0301	0.0254	
Nu_{ϵ_u}	14.93	14.62	14.28	14.03	13.78	12.93	11.96	11.61	11.76	13.69	17.50	16.04	16.07	16.02	16.06	16.07	16.07	16.04	16.01	15.97	15.97	15.58	15.04	14.47	13.85	13.42	13.09	12.85	12.63	12.47	12.13	
$N u_{\epsilon_{\theta}}$	14.91	14.62	14.27	14.03	13.77	12.93	11.98	11.62	11.77	13.74	17.44	16.03	16.07	16.02	16.06	16.05	16.06	16.03	16.00	15.97	15.95	15.55	15.04	14.47	13.85	13.41	13.09	12.84	12.62	12.47	12.16	
Nu_w	14.93	14.62	14.27	14.02	13.78	12.93	11.97	11.62	11.77	13.74	17.44	16.03	16.07	16.03	16.06	16.05	16.07	16.03	16.00	15.98	15.95	15.56	15.04	14.47	13.85	13.41	13.09	12.84	12.62	12.47	12.17	
Re_S	009	700	800	900	1000	1414	2236	3162	4472	7071	$10\ 000$	10	20	30	40	50	60	70	80	90	100	200	300	400	500	600	700	800	006	1000	1200	
Pr	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	3.0	3.0	3.0	3.0	3.0	3.0	3.0	3.0	3.0	3.0	3.0	3.0	3.0	3.0	3.0	3.0	3.0	3.0	3.0	3.0	
Ra	10^7	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^7	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	oage.
N_{BL}	29	29	30	30	30	31	33	33	33	30	27	28	28	28	28	28	28	28	28	28	28	28	29	30	30	31	31	31	32	32	32	ee next j
+22	0.64	0.71	0.77	0.82	0.88	1.07	1.37	1.64	1.98	2.69	3.62	0.06	0.09	0.11	0.12	0.15	0.16	0.18	0.19	0.20	0.21	0.30	0.37	0.43	0.51	0.57	0.62	0.67	0.71	0.75	0.83	caption s
$\frac{i}{2}$	0.01	0.01	0.01	0.01	0.01	0.01	0.02	0.02	0.02	0.03	0.04	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	3. For
$\downarrow^{\mathcal{V}}$	0.55	0.61	0.66	0.70	0.75	0.91	1.17	1.39	1.69	2.30	3.08	0.05	0.07	0.09	0.10	0.13	0.13	0.15	0.16	0.17	0.18	0.26	0.31	0.37	0.43	0.49	0.53	0.57	0.60	0.64	0.70	Table
$^+\!x$	0.55	0.61	0.66	0.70	0.75	0.91	1.17	1.39	1.69	2.30	3.08	0.05	0.07	0.09	0.10	0.13	0.13	0.15	0.16	0.17	0.18	0.26	0.31	0.37	0.43	0.49	0.53	0.57	0.60	0.64	0.70	
N^2	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	
N_y	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	
N_x	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	
Γ_y	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	
Γ_x	8	8	8	8	8	8	8	8	8	8	×	8	8	8	8	8	8	8	8	8	8	×	×	8	8	8	8	8	8	8	8	
System	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	

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u_{ϵ_u} C_S	2.06 0.0204	2.80 0.016/	2.58 0.0106	1.04 0.0087	5.97 0.0078 ^t	3.18 0.00735	0.0070	2.71 0.0068	1.00 0.0062	.93 0.0060 ⁻	5.13 2.13	5.15 1.03	6.14 0.553	5.11 0.481	5.08 0.368	5.03 0.324	5.02 0.264	5.91 0.253	5.97 0.217	5.85 0.201	5.10 0.108	1.34 0.0751	3.78 0.0552	3.18 0.0492	3.00 0.0411	2.77 0.0354	2.87 0.0332	2.89 0.0303	3.04 0.028	2.21 0.0211	2.19 0.017	
$N u_{\epsilon_{ heta}}$ $N u$	12.06 12	12.81	12.59 12	14.00 14	15.97 15	18.16 18	20.31 20	22.74 22	24.11 24	26.36 26	16.13 16	16.14 16	16.12 16	16.11 16	16.09 16	16.04 16	16.02 16	15.93 15	15.98 15	15.85 15	15.09 15	14.33 14	13.78 13	13.18 13	12.91 13	12.67 12	12.84 12	12.94 12	13.07 13	12.20 12	12.19 12	
Nu_{w}	12.05	12.81	12.59	14.00	15.98	18.15	20.34	22.73	24.18	26.23	16.13	16.14	16.13	16.11	16.09	16.04	16.02	15.93	15.98	15.85	15.09	14.33	13.78	13.18	12.91	12.67	12.84	12.94	13.07	12.20	12.19	
Re_S	1500	2000	3000	4000	5000	6000	7000	8000	0006	$10\ 000$	10	20	30	40	50	60	70	80	90	100	200	300	400	500	009	700	800	900	1000	1200	1500	
Pr	3.0	3.U	3.0	3.0	3.0	3.0	3.0	3.0	3.0	3.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	
Ra	10^{7}	, 01	10	10'	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^7	10^7	10^{7}	10^{7}	10^7	10^{7}	10^{7}	10^7	10^7	10^7	10^7	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^7	10^{7}	10^{7}	10^{7}	10^{7}	10^7	
N_{BL}	33	27 77	32	30	28	26	25	23	23	22	41	41	41	41	41	41	41	41	41	41	42	44	45	46	46	47	46	46	46	48	48	
2 ^m / ₂	0.92	71.1	1.34	1.62	1.92	2.23	2.54	2.86	3.08	3.37	0.04	0.06	0.06	0.08	0.09	0.10	0.10	0.11	0.12	0.13	0.19	0.23	0.27	0.32	0.35	0.38	0.42	0.45	0.48	0.50	0.56	
$+_{w}^{Z}$	0.01	10.0	0.02	0.02	0.02	0.03	0.03	0.03	0.04	0.04	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	
y^+	0.79	cy.U	1.14	1.38	1.64	1.90	2.16	2.44	2.63	2.87	0.04	0.06	0.06	0.08	0.08	0.09	0.10	0.11	0.12	0.12	0.18	0.23	0.26	0.31	0.34	0.36	0.40	0.43	0.46	0.48	0.54	
x^+	0.79	cy.U	1.14	1.38	1.64	1.90	2.16	2.44	2.63	2.87	0.04	0.06	0.06	0.08	0.08	0.09	0.10	0.11	0.12	0.12	0.18	0.23	0.26	0.31	0.34	0.36	0.40	0.43	0.46	0.48	0.54	
N_z	256	007	256	256	256	256	256	256	256	256	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	
N_y	768	/08	768	768	768	768	768	768	768	768	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	
N_x	1536	0501	1536	1536	1536	1536	1536	1536	1536	1536	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	
Γ_y	4 •	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	
Γ_x	~ ~	×	×	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	
System	CRB	CKB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	

Transport in sheared Rayleigh-Bénard convection

C_S	7 0.0128	0.00942	0.00818	0.0075	3 0.007	0.0065	0.00637	3 0.0061	0.00605	0.893	0.281	1 0.188	7 0.183	0.138	0.124	1 0.105	0.0879	0.0836	5 0.0743	0.0587	5 0.056	7 0.0481	0.0426	0.0392	0.0371	3 0.0356	0.0318	7 0.0292	7 0.0292	0.0284	3 0.0265	
$N u_{\epsilon_u}$	12.27	13.96	16.39	19.14	21.88	24.31	27.10	29.53	32.42	30.85	30.78	30.84	30.87	30.81	30.74	30.64	30.71	30.49	30.55	30.50	29.66	29.77	29.52	29.20	28.67	27.48	28.11	28.07	27.87	27.67	27.48	
$N u_{\epsilon_{ heta}}$	12.28	13.98	16.44	19.13	21.72	24.24	27.08	29.60	32.54	30.86	30.68	30.80	30.80	30.79	30.73	30.70	30.65	30.54	30.53	30.45	30.10	29.80	29.53	29.20	28.74	28.57	28.16	28.07	27.83	27.68	27.42	
Nu_{W}	12.28	13.98	16.44	19.14	21.72	24.23	27.10	29.63	32.59	30.85	30.68	30.80	30.81	30.79	30.72	30.69	30.66	30.55	30.49	30.45	29.97	29.80	29.54	29.19	28.73	28.46	28.16	28.07	27.82	27.68	27.40	
Re_S	2000	3000	4000	5000	6000	7000	8000	0006	$10\ 000$	100	200	300	400	500	009	700	800	900	1000	1200	1400	1600	1800	2000	2200	2400	2600	2800	2900	3000	3200	
Pr	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	
Ra	10^7	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	
N_{BL}	47	44	40	37	35	33	31	29	28	29	29	29	29	29	29	29	29	29	29	29	29	29	29	30	30	30	30	30	30	30	31	
$+\frac{\pi}{2}$	0.64	0.83	1.03	1.24	1.43	1.61	1.82	2.01	2.22	0.27	0.30	0.37	0.49	0.53	0.60	0.65	0.68	0.74	0.78	0.83	0.95	1.00	1.06	1.13	1.21	1.29	1.32	1.37	1.42	1.44	1.49	a control
$z_{\chi}^{+,\mu}$	0.02	0.02	0.03	0.03	0.03	0.04	0.04	0.05	0.05	0.01	0.01	0.01	0.01	0.01	0.01	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.03	0.03	0.03	0.03	0.03	0.03	0.03	0.03	0.04	با م
γ_+	0.62	0.80	1.00	1.20	1.39	1.56	1.76	1.94	2.15	0.26	0.29	0.36	0.47	0.51	0.58	0.63	0.66	0.72	0.75	0.80	0.92	0.97	1.03	1.09	1.17	1.25	1.28	1.32	1.37	1.40	1.44	E
^{+}x	0.62	0.80	1.00	1.20	1.39	1.56	1.76	1.94	2.15	0.26	0.29	0.36	0.47	0.51	0.58	0.63	0.66	0.72	0.75	0.80	0.92	0.97	1.03	1.09	1.17	1.25	1.28	1.32	1.37	1.40	1.44	
N_z	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	384	
N_y	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	1024	
N_x	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	2048	
Γ_y	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	
Γ_x	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	×	8	8	8	8	8	8	
System	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB									

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C_S	0.0252	0.0232	0.022	0.0179	0.0153	0.0134	0.0121	0.011	0.0101	2.15	1.15	0.644	0.524	0.417	0.358	0.294	0.259	0.237	0.216	0.11	0.0748	0.0595	0.0486	0.0414	0.0361	0.0322	0.0293	0.0266	0.0227	0.0211	
Nu_{ϵ_u}	27.02 27.01	26.74	26.58	26.18	25.77	25.19	25.36	25.61	25.64	8.25	8.31	8.31	8.28	8.29	8.26	8.25	8.19	8.21	8.17	7.87	7.36	7.04	6.74	6.56	6.44	6.38	6.35	6.32	6.30	7.06	
$Nu_{\epsilon_{\theta}}$	27.16 27.00	26.80	26.63	26.07	25.60	25.19	25.33	25.49	25.68	8.25	8.31	8.31	8.28	8.29	8.26	8.25	8.19	8.21	8.16	7.87	7.36	7.04	6.74	6.56	6.43	6.38	6.35	6.31	6.30	7.06	
Nu_w	27.16 26.00	26.80	26.60	26.05	25.57	25.19	25.32	25.49	25.69	8.25	8.31	8.30	8.28	8.29	8.26	8.25	8.19	8.21	8.17	7.86	7.37	7.04	6.74	6.56	6.43	6.39	6.35	6.32	6.31	7.07	
Re_S	3400 3600	3800	4000	5000	6000	7000	8000	0006	10 000	10	20	30	40	50	60	70	80	90	100	200	300	400	500	600	700	800	006	1000	1200	1500	
Pr	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	
Ra	10 ⁸	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	age.
N_{BL}	31	5 55	31	32	32	32	32	32	32	40	40	40	40	40	40	40	40	40	40	41	42	43	44	45	45	45	46	46	46	43	e next p
z_{m+z}^{+z}	1.54	1.65	1.69	1.91	2.12	2.32	2.52	2.69	2.86	0.06	0.09	0.10	0.13	0.14	0.16	0.16	0.18	0.19	0.20	0.29	0.35	0.42	0.48	0.53	0.57	0.62	0.67	0.70	0.78	0.94	uption se
$+\frac{2}{3}$.04	.04	.04	0.05	0.05).06	.06	.07	.07	00.0	00.0	00.0	00.0	00.0	00.0	00.0	00.0	00.0	00.0	00.0	00.0	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	For ca
\mathbf{v}^+	.49	5 <u>9</u>	.6	.85 (.05 (.24 (.44	.60	. 17	.05 (.08) 60.	.11 (.12 (.13 (.14 (.15 (.16 (.17 (.24 (.30 (.36 (.41	.45	.49	.53	.57	09.	.67	.80	Table 3
+	49 1	5 9	64 1	85 1	05 2	24 2	44	60 2	77 2	05 0	08 0	0 60	11 0	12 0	.13 C	.14	15 0	.16 C	17 0	24 0	30 0	36 0	41 C	45 0	49 0	53 0	57 0	60 0	67 0	80 0	
x	-				<i>с</i> і	<i>с</i> і	7	<i>с</i> і	6.	0	0	0.	0	0	0	0	0	0	0	0.	0.	0.	Ö	0.	0.	0.	0	0.	0	0	
N_z	384	384	384	384	384	384	384	384	384	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	
N_y	1024	1024	1024	1024	1024	1024	1024	1024	1024	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	
N_x	2048 2048	2048	2048	2048	2048	2048	2048	2048	2048	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	
Γ_y	4 ~	+ 4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	
Γ_x	× •	∞	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	
System	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	CRB	PRB																					

Transport in sheared Rayleigh-Bénard convection

C_S	0.0166	0.0104	0.00866	0.0083	0.00811	0.00785	0.00755	0.00731	0.00707	4.76	2.57	1.56	1.07	0.923	0.82	0.708	0.607	0.588	0.511	0.249	0.159	0.121	0.0974	0.0799	0.0695	0.06	0.0537	0.0487	0.0415	0.0338	0.0262	
Nu_{ϵ_u}	7.33	6.64	7.55	8.49	10.08	11.05	11.60	11.60	12.25	15.27	15.30	15.27	15.25	15.28	15.27	15.30	15.23	15.26	15.30	15.21	15.22	15.13	15.09	15.03	14.92	14.81	14.67	14.60	14.31	13.86	13.21	
$N u_{\epsilon_{ heta}}$	7.32	6.65	7.57	8.45	9.99	11.03	11.54	11.55	12.16	15.27	15.30	15.27	15.25	15.29	15.26	15.29	15.23	15.27	15.31	15.21	15.21	15.12	15.10	15.02	14.91	14.81	14.67	14.60	14.31	13.85	13.20	
Nu_{w}	7.34	6.67	7.59	8.50	10.05	11.10	11.64	11.68	12.35	15.27	15.29	15.27	15.24	15.28	15.26	15.29	15.23	15.26	15.30	15.20	15.21	15.12	15.09	15.02	14.91	14.82	14.67	14.59	14.31	13.85	13.20	
Re_S	2000	3000	4000	5000	0009	7000	8000	0006	10 000	10	20	30	40	50	60	70	80	90	100	200	300	400	500	600	700	800	006	1000	1200	1500	2000	
Pr	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	0.5	
Ra	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^{6}	10^7	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	10^{7}	page.
N_{BL}	42	44	41	39	36	34	33	33	32	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	30	30	31	ee next j
+ ² ²	1.11	1.32	1.61	1.97	2.34	2.68	3.01	3.32	3.63	0.09	0.14	0.16	0.18	0.21	0.23	0.25	0.27	0.30	0.31	0.43	0.52	0.60	0.67	0.73	0.80	0.85	0.90	0.95	1.06	1.19	1.40	caption s
z^{+s}	0.01	0.02	0.02	0.02	0.03	0.03	0.04	0.04	0.04	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.02	3. For e
γ_+	0.95	1.13	1.37	1.68	1.99	2.28	2.56	2.83	3.10	0.08	0.12	0.14	0.15	0.18	0.20	0.22	0.23	0.25	0.26	0.37	0.44	0.51	0.57	0.62	0.68	0.72	0.77	0.81	0.90	1.02	1.19	Table
x^+	0.95	1.13	1.37	1.68	1.99	2.28	2.56	2.83	3.10	0.08	0.12	0.14	0.15	0.18	0.20	0.22	0.23	0.25	0.26	0.37	0.44	0.51	0.57	0.62	0.68	0.72	0.77	0.81	0.90	1.02	1.19	
N_z	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	
N_y	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	
N_x	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	
Γ_y	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	
Γ_x	8	8	8	8	8	8	8	8	×	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	
System	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB									

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C_S	0.0183 0.014	0.0114	0.00981	0.00878	0.00812	0.00781	0.00743	3.95	1.73	0.77	0.395	0.182	0.123	0.0913	0.0746	0.0625	0.0548	0.0488	0.0442	0.0404	0.03	0.0197	0.0144	0.0108	0.0082	0.00735	2.14	1.23	0.829	0.636	
Nu_{ϵ_u}	12.24 11.78	11.67	11.68	11.79	12.09	12.58	13.16	15.74	15.74	15.73	15.75	15.68	15.55	15.39	15.14	14.97	14.71	14.50	14.21	14.00	13.18	12.20	11.87	11.91	13.30	17.04	16.12	16.13	16.12	16.11	
$Nu_{\epsilon_{\theta}}$	12.23 11.77	11.67	11.68	11.80	12.08	12.57	13.14	15.74	15.74	15.73	15.75	15.67	15.55	15.39	15.15	14.97	14.71	14.51	14.21	13.99	13.17	12.18	11.85	11.90	13.25	17.02	16.11	16.12	16.11	16.10	
Nu_{W}	12.24 11.77	11.68	11.69	11.82	12.10	12.59	13.16	15.74	15.73	15.73	15.74	15.67	15.54	15.39	15.15	14.97	14.71	14.51	14.21	13.99	13.17	12.18	11.86	11.91	13.29	17.11	16.11	16.12	16.12	16.10	
Re_S	3000 4000	5000	0009	7000	8000	0006	10000	10	20	50	100	200	300	400	500	600	700	800	006	1000	1414	2236	3162	4472	7071	10000	10	20	30	40	
Pr	0.5 0.5	0.5	0.5	0.5	0.5	0.5	0.5	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	3.0	3.0	3.0	3.0	
Ra	10^{7}	10^7	10^7	10^7	10^7	10^7	10^{7}	10^7	10^7	10^7	10^7	10^7	10^7	10^7	10^7	10^7	10^7	10^{7}	10^7	10^7	10^7	10^7	10^7	10^7	10^7	10^{7}	10^7	10^7	10^7	10^7	oage.
N_{BL}	33 33	33	33	33	32	32	31	28	28	28	28	28	29	29	29	29	29	30	30	30	31	32	33	33	31	27	28	28	28	28	ee next j
+ ² m	1.75 2.05	2.30	2.57	2.84	3.12	3.44	3.73	0.09	0.11	0.19	0.27	0.37	0.45	0.52	0.59	0.65	0.71	0.76	0.82	0.87	1.06	1.36	1.64	2.01	2.77	3.71	0.06	0.10	0.12	0.14	caption se
z_w^{+2}	0.02 0.03	0.03	0.03	0.03	0.04	0.04	0.05	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.02	0.02	0.02	0.03	0.05	0.00	0.00	0.00	0.00	3. For (
y^+	1.49 1.74	1.96	2.19	2.42	2.66	2.93	3.17	0.07	0.10	0.16	0.23	0.31	0.39	0.45	0.50	0.55	0.60	0.65	0.70	0.74	0.90	1.16	1.40	1.71	2.36	3.16	0.05	0.08	0.10	0.12	Table
x^+	1.49 1.74	1.96	2.19	2.42	2.66	2.93	3.17	0.07	0.10	0.16	0.23	0.31	0.39	0.45	0.50	0.55	0.60	0.65	0.70	0.74	0.90	1.16	1.40	1.71	2.36	3.16	0.05	0.08	0.10	0.12	
N_z	256 256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	
N_y	768 768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	
N_x	1536 1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	
Γ_y	44	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	
Γ_x	~ ~	×	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	×	
System	PRB PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	

Transport in sheared Rayleigh-Bénard convection

System	Γ_x	Γ_y	N_x	N_y	N_z	x^+	y_{+}	$+\frac{2}{3}$	z_m^{z+z}	N_{BL}	Ra	Pr	Re_S	Nu_{w}	$Nu_{\epsilon_{ heta}}$	Nu_{ϵ_u}	C_S
PRB	8	4	1536	768	256	0.13	0.13	0.00	0.15	28	10^7	3.0	50	16.05	16.05	16.06	0.488
PRB	8	4	1536	768	256	0.15	0.15	0.00	0.17	28	10^7	3.0	09	16.02	16.02	16.02	0.452
PRB	8	4	1536	768	256	0.15	0.15	0.00	0.18	28	10^7	3.0	70	16.05	16.05	16.05	0.347
PRB	8	4	1536	768	256	0.16	0.16	0.00	0.19	28	10^7	3.0	80	16.03	16.03	16.04	0.308
PRB	8	4	1536	768	256	0.17	0.17	0.00	0.20	28	10^7	3.0	90	16.00	16.00	16.01	0.275
PRB	8	4	1536	768	256	0.18	0.18	0.00	0.22	28	10^7	3.0	100	15.98	15.98	15.99	0.25
PRB	8	4	1536	768	256	0.26	0.26	0.00	0.31	28	10^7	3.0	200	15.55	15.55	15.56	0.126
PRB	8	4	1536	768	256	0.33	0.33	0.00	0.39	29	10^7	3.0	300	15.02	15.01	15.02	0.0897
PRB	8	4	1536	768	256	0.39	0.39	0.01	0.46	30	10^7	3.0	400	14.43	14.43	14.44	0.0705
PRB	8	4	1536	768	256	0.45	0.45	0.01	0.52	30	10^7	3.0	500	13.97	13.97	13.98	0.0585
PRB	8	4	1536	768	256	0.49	0.49	0.01	0.58	31	10^7	3.0	600	13.61	13.60	13.62	0.0497
PRB	8	4	1536	768	256	0.54	0.54	0.01	0.63	31	10^7	3.0	700	13.28	13.28	13.28	0.0436
PRB	8	4	1536	768	256	0.58	0.58	0.01	0.68	31	10^7	3.0	800	13.01	13.01	13.02	0.0384
PRB	8	4	1536	768	256	0.62	0.62	0.01	0.72	32	10^7	3.0	006	12.78	12.78	12.80	0.0345
PRB	8	4	1536	768	256	0.65	0.65	0.01	0.76	32	10^7	3.0	1000	12.65	12.64	12.65	0.0313
PRB	8	4	1536	768	256	0.72	0.72	0.01	0.85	32	10^{7}	3.0	1200	12.39	12.39	12.42	0.0266
PRB	8	4	1536	768	256	0.81	0.81	0.01	0.95	32	10^7	3.0	1500	12.26	12.25	12.27	0.0217
PRB	8	4	1536	768	256	0.95	0.95	0.01	1.12	32	10^7	3.0	2000	12.20	12.19	12.20	0.0167
PRB	8	4	1536	768	256	1.20	1.20	0.02	1.41	33	10^{7}	3.0	3000	12.09	12.08	12.08	0.0119
PRB	8	4	1536	768	256	1.44	1.44	0.02	1.69	31	10^7	3.0	4000	13.18	13.14	13.27	0.00954
PRB	8	4	1536	768	256	1.71	1.71	0.02	2.01	29	10^7	3.0	5000	14.92	14.85	14.90	0.00865
PRB	8	4	1536	768	256	1.99	1.99	0.03	2.33	27	10^7	3.0	0009	17.68	17.57	17.60	0.00808
PRB	8	4	1536	768	256	2.30	2.30	0.03	2.69	25	10^7	3.0	7000	20.20	20.02	20.23	0.00793
PRB	×	4	1536	768	256	2.58	2.58	0.04	3.03	23	10^{7}	3.0	8000	22.88	22.66	22.89	0.00768
PRB	×	4	1536	768	256	2.86	2.86	0.04	3.35	22	10^7	3.0	0006	25.00	24.70	25.00	0.00743
PRB	×	4	1536	768	256	3.13	3.13	0.04	3.67	21	10^7	3.0	10000	26.99	26.59	26.96	0.0072
PRB	8	4	1536	768	256	0.05	0.05	0.00	0.06	28	10^7	5.0	10	16.20	16.19	16.20	2.16
PRB	8	4	1536	768	256	0.08	0.08	0.00	0.09	28	10^7	5.0	20	16.13	16.12	16.13	1.2
PRB	8	4	1536	768	256	0.10	0.10	0.00	0.11	28	10^7	5.0	30	16.11	16.11	16.11	0.778
PRB	8	4	1536	768	256	0.11	0.11	0.00	0.13	28	10^7	5.0	40	16.10	16.10	16.10	0.593
PRB	8	4	1536	768	256	0.12	0.12	0.00	0.14	28	10^7	5.0	50	16.10	16.10	16.11	0.42
							Table	e 3. For	caption s	ee next p	age.						

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C_S).358	1.241 268	0.239	0.218	0.118	0.0838	0.0648	0.0528	0.0446	0.0387	0.0343	0.0307	0.0278	0.0234	0.0203	0.0151	0.0107	0.00912	0.00858	0.00822	16700.0	0.00759	0.00733	.0071	2.17	5.35	2.55	2.16	.89	.19	1.22	
n	80	70 8	9. S	92 (18	38 () 62	30 (00	80	61	54 (4	34 (46 () 66	47 (09	55 (94 (41 (97 (95 (49 (76	97 (86	2	90	83	86	
Nu_{ϵ}	16. 16	10.	15.	15.	15.	14.	13.	13.	13.	12.	12.	12.	12.	12.	12.	11.	13.	15.	18.	21.	24.	26.	28.	30.	30.	30.	30.	31.	30.	30.	30.	
$N u_{\epsilon_{\theta}}$	16.08 16.01	10.01 16.02	15.94	15.91	15.17	14.37	13.78	13.30	12.99	12.78	12.59	12.52	12.43	12.34	12.48	11.99	13.45	15.55	18.32	21.52	24.14	26.39	27.98	29.42	30.75	30.86	30.78	30.98	30.92	30.78	30.87	
łw	80.	10.	946:	.91	.18	.37	.78	.30	66.	.78	.59	.53	.43	.34	.49	.01	.48	.62	.47	.72	.50	.90	.72	.42	.73	.85	.78	.98	.92	.76	.87	
N_l	16	01 Y	12	15	15	14	13	13	12	12	12	12	12	12	12	12	13	15	18	21	24	26	28	30	30	30	30	30	30	30	30	
Re_S	09	0/ 08	06 06	100	200	300	400	500	009	700	800	900	1000	1200	1500	2000	3000	4000	5000	6000	7000	8000	0006	$10\ 000$	10	20	30	40	50	60	70	
Pr	5.0	0.0 2	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	5.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	
Ra	10^{7}	107	10^{7}	10^7	10^7	10^7	10^7	10^7	10^{7}	10^7	10^7	10^7	10^7	10^7	10^7	10^7	10^7	10^{7}	10^{7}	10^7	10^7	10^7	10^7	10^{7}	10^8	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	10^{8}	page.
N_{BL}	28 28	0 X 7 C	28 78	28	29	30	30	31	31	32	32	32	32	32	32	33	31	28	26	24	22	21	21	20	30	30	30	30	30	30	30	ee next j
z_m^+	0.16	0.18	0.19	0.20	0.30	0.38	0.44	0.50	0.55	0.60	0.64	0.68	0.72	0.79	0.92	1.06	1.34	1.65	2.00	2.35	2.69	3.01	3.33	3.64	0.04	0.15	0.14	0.17	0.20	0.19	0.22	caption s
$z_w^{z^+}$	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.02	0.02	0.02	0.03	0.03	0.04	0.04	0.04	0.00	0.00	0.00	0.00	0.00	0.00	0.00	e 3. For
y^+	0.13	0.15 0	0.16	0.17	0.25	0.32	0.37	0.42	0.47	0.51	0.55	0.58	0.61	0.68	0.79	0.91	1.14	1.41	1.71	2.00	2.29	2.57	2.84	3.10	0.04	0.14	0.13	0.16	0.19	0.18	0.21	Table
x^+	0.13	CT.0	0.16	0.17	0.25	0.32	0.37	0.42	0.47	0.51	0.55	0.58	0.61	0.68	0.79	0.91	1.14	1.41	1.71	2.00	2.29	2.57	2.84	3.10	0.04	0.14	0.13	0.16	0.19	0.18	0.21	
N_z	256 256	956	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	256	384	384	384	384	384	384	384	
N_y	768 76 º	00/ 76.8	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	768	1024	1024	1024	1024	1024	1024	1024	
N_x	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	1536	2048	2048	2048	2048	2048	2048	2048	
Γ_y	4 -	4 4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	
Γ_x	~ ~	0 x	⊳ ∞	8	8	×	×	×	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	×	×	×	8	
System	PRB	PRR	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	PRB	

Transport in sheared Rayleigh-Bénard convection

ystem	Γ_x	Γ_{y}	N_x	N_y	N_z	^+x	\mathbf{y}^+	z_w^{+2}	τ ² +	N_{BL}	Ra	Pr	Re_S	Nu_w	$N u_{\epsilon_{\theta}}$	Nu_{ϵ_u}	C_S
Ð	8	4	2048	1024	384	0.23	0.23	0.00	0.24	30	10^{8}	1.0	80	30.72	30.72	30.77	1.11
æ	8	4	2048	1024	384	0.25	0.25	0.00	0.27	30	10^{8}	1.0	90	30.69	30.70	30.73	1.05
ß	8	4	2048	1024	384	0.21	0.21	0.00	0.22	30	10^{8}	1.0	100	30.93	30.92	30.89	0.562
RB	8	4	2048	1024	384	0.32	0.32	0.00	0.33	30	10^{8}	1.0	200	30.85	30.85	30.83	0.329
RB	8	4	2048	1024	384	0.42	0.42	0.00	0.44	30	10^{8}	1.0	300	30.81	30.81	30.84	0.254
RB	8	4	2048	1024	384	0.47	0.47	0.00	0.49	30	10^{8}	1.0	400	30.69	30.68	30.61	0.184
RB	8	4	2048	1024	384	0.53	0.53	0.00	0.55	30	10^{8}	1.0	500	30.78	30.80	30.78	0.147
RB	8	4	2048	1024	384	0.58	0.58	0.00	0.61	30	10^{8}	1.0	600	30.75	30.75	30.82	0.123
PRB	8	4	2048	1024	384	0.63	0.63	0.01	0.66	30	10^{8}	1.0	700	30.72	30.73	30.69	0.107
PRB	8	4	2048	1024	384	0.68	0.68	0.01	0.71	30	10^{8}	1.0	800	30.58	30.59	30.57	0.0935
PRB	8	4	2048	1024	384	0.70	0.70	0.01	0.74	30	10^{8}	1.0	006	30.45	30.45	30.43	0.0804
PRB	8	4	2048	1024	384	0.75	0.75	0.01	0.78	31	10^{8}	1.0	1000	30.34	30.35	30.29	0.0731
PRB	8	4	2048	1024	384	0.85	0.85	0.01	0.89	31	10^{8}	1.0	1200	30.21	30.22	30.23	0.066
PRB	8	4	2048	1024	384	0.96	0.96	0.01	1.00	31	10^{8}	1.0	1500	30.01	30.01	29.95	0.0537
PRB	8	4	2048	1024	384	1.08	1.08	0.01	1.12	31	10^{8}	1.0	2000	29.33	29.32	29.38	0.0379
PRB	8	4	2048	1024	384	1.36	1.36	0.01	1.42	32	10^{8}	1.0	3000	27.95	27.96	27.96	0.027
PRB	8	4	2048	1024	384	1.60	1.60	0.01	1.67	33	10^{8}	1.0	4000	26.93	26.93	26.90	0.0211
PRB	8	4	2048	1024	384	1.81	1.81	0.02	1.89	33	10^{8}	1.0	5000	26.17	26.17	26.26	0.0172
PRB	8	4	2048	1024	384	2.01	2.01	0.02	2.10	33	10^{8}	1.0	0009	25.62	25.59	25.60	0.0147
PRB	8	4	2048	1024	384	2.18	2.18	0.02	2.28	34	10^{8}	1.0	7000	25.35	25.34	25.45	0.0128
PRB	8	4	2048	1024	384	2.34	2.34	0.02	2.44	34	10^{8}	1.0	8000	25.29	25.27	25.37	0.0112
PRB	8	4	2048	1024	384	2.51	2.51	0.02	2.62	34	10^{8}	1.0	0006	25.30	25.27	25.34	0.0102
PRB	×	4	2048	1024	384	2.67	2.67	0.02	2.78	33	10^{8}	1.0	10000	25.52	25.49	25.52	0.00931
able 3. 5	Simulat	ions co	nsidered i	in this worl	k. The asp	sect ratios	of the c	omputatic	onal dom	ain are g	iven by	Γ_x in th	e streamwi	se direction	and Γ_y in t	he spanwis	e direction
he values	of N_x ,	N_y and	$1 N_z$ indic	tate the nui	mber of g	rid points	s in the s	treamwis	e, spanwi	se and w	/all-norr	nal dire	ctions. The	e grid spacir	ng in wall u	nits in the	streamwise
nd spanw	ise dire	ctions	is given t	by Δx^+ and	d ∆y ⁺ , r€	espectivel	y. The w	all-norm	al grid sţ	acing in	wall un	uits at th	ne wall and	l the mid-he	sight are giv	ven by Δz_i	r_{c}^{+} and Δz_{c}^{+}
	E T	haire	. of and .	soints in th	on llow o	and dire	inter min	+ hin the +	T Lound	T is sin	h N7	How.	Mr. indi	are the second	June of Miner	altanna tha	" optimized

respectively. The number of grid points in the wall-normal direction within the thermal BL is given by N_{BL} . Here, Nu_w indicates the value of Nusselt number computed using the average gradient of reduced temperature profile at the walls using (1.5) and Nu_{e_0} and Nu_{e_0} indicate the Nusselt numbers that satisfy the global balance of thermal and kinetic dissipation (2.1) and (2.2).

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