

## Origin and Distribution of Water Amongst the Inner Planets

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**Abstract.** A new model for the origin and bulk chemical composition of the inner planets, especially for their water content, is reported.

There is a growing body of evidence that the planets of the inner Solar system condensed within narrow, compositionally-distinct annuli, close to their present orbital radii (Drake & Righter 2002). Such a picture is consistent with the Laplacian nebular hypothesis, namely that the planetary system had formed from a concentric family of gas rings. Such rings were shed by the contracting proto-solar cloud [hereafter PSC] as a means for disposing of excess spin angular momentum (Prentice 1978). A new model for the PSC has been constructed. It consists of an adiabatic convective core surrounded by a super-adiabatic envelope of negative polytropic index. This structure is suggested by numerical simulations of supersonic thermal convection in a model atmospheric layer (Prentice & Dyt 2003). The cloud possesses a radial turbulent stress whose ratio to the gas pressure achieves a maximum value  $\sim 15$  at the core/envelope boundary. If the controlling parameters stay constant, the PSC contracts homologously and sheds gas rings whose mean orbital radii  $R_n$  ( $n = 0, 1, 2, \dots$ ) are nearly geometrically spaced. The ring mean temperatures  $T_n$  vary with  $R_n$  as  $T_n \simeq A/R_n^{0.9}$ , where  $A$  is a constant (Prentice 2001a). Choosing iron-rich Mercury to calibrate  $A$  (Prentice 2001b), Venus forms at 917 K and is totally anhydrous. The initial Earth (678 K) has 0.0023, by mass fraction, of water tied up in tremolite. Mars (460 K) contains an  $\text{H}_2\text{O}$  mass fraction of 0.00295 in tremolite and 0.0027 in (Na,K)OH. The asteroids (275 K) contain 0.0027 of  $\text{H}_2\text{O}$  in (Na,K)OH. Mars is thus the most water-rich of all the inner planets. Other predicted bulk constituents of Mars are:  $\text{MgAl}_2\text{O}_4$  (0.0324),  $\text{MgSiO}_3$ - $\text{Mg}_2\text{SiO}_4$  (0.3677),  $\text{Fe}_2\text{SiO}_4$  (0.1796), Fe-Ni-Cr (0.0533), (Fe-Ni)S (0.2042), MnS & ZnS (0.0050), NaCl (0.0016),  $\text{Cr}_2\text{O}_3$  &  $\text{FeTiO}_3$  (0.0065) and  $\text{P}_2\text{O}_5$  (0.0031).

## References

- Drake, M. J., & Righter, K. 2002, *Nature*, 416, 39  
Prentice, A. J. R. 1978, *Moon & Planets*, 19, 341  
Prentice, A. J. R. 2001a, *Earth Moon & Planets*, 87, 11  
Prentice, A. J. R. 2001b in *LPI Contribution No. 1097, Workshop on Mercury: Space Environment, Surface, and Interior*, ed. M. S. Robinson & G. J. Taylor (Houston: Lunar & Planetary Institute), 81  
Prentice, A. J. R., & Dyt, C. P. 2003, *MNRAS*, 341, 644