Ultraviolet Emission from Star-formation in Selected Gas-rich Early-type Galaxies

Lerothodi L. Leeuw^{1,2}

¹College of Graduate Studies, UNISA, P. O. Box 392, UNISA Pretoria, 0003, South Africa
²SETI Institute, 189 Bernardo Avenue, Suite 100, Mountain View, CA 94043, USA
email: lerothodi@alum.mit.edu

Abstract. We present GALEX ultraviolet (UV) emission results of star-formation in a small sample of nearby, gas-rich early-type galaxies. The first observational evidence of star-formation in this sample was presented by Leeuw et~al.~2008, using $350~\mu m$ continuum data. The measured far-infrared (far-IR) excess of these galaxies showed that the most likely and dominant heating source of the observed $350~\mu m$ continuum emission from dust is star-formation, that could have been triggered by an accretion or merger event. Consistent with starbursts that are less than 1 Gyr (e.g., Kaviraj 2010), the GALEX near-UV (NUV) minus SDSS r-band emission of the galaxies is < 5.5. The UV results corroborate those of mid-IR to radio data for the sample.

 $\label{eq:Keywords.} \textbf{Keywords.} \ \ \text{galaxies: elliptical and lenticular, cD-galaxies: evolution-galaxies: ISM-infrared: general-submillimeter-ultraviolet: general-submillimeter-s$

Synopsis

We summarize ultraviolet (UV) emission results of on-going star-formation determined from archival GALEX and SDSS data for a small sample of nearby early-type galaxies, that were known to have CO in centrally located gas disks. Like classic ellipticals, the luminosity profiles of these galaxies follow the de Vaucoleur, $r^{1/4}$ law. However, they represent a spread of merger tracers or ages, from galaxies that have been morphologicaly classified as on-going or early-age major mergers to very-late accretion or quiescent systems (Leeuw et al. 2008; 2011). Using 350 µm continuum data, Leeuw et al. 2008 presented the first observational evidence of star-formation in this sample. The measured far-infrared (far-IR) excess of these galaxies showed that the most likely and dominant heating source of the observed $350\,\mu\mathrm{m}$ continuum emission from dust is star-formation, that could have been triggered by an accretion or merger event and is stronger in the "most recent mergers". Consistent with starbursts that are less than 1 Gyr (e.g., Kaviraj 2010), the GALEX near-UV (NUV) minus SDSS r-band emission of the galaxies is < 5.5. Further, the sample dwarf elliptical NGC 4476 and intermediate-age elliptical NGC 5666, that respectively have only a hint and a dense ring of star-formation, based on their far-IR data, have NUV-r = 4.58 and 2.64. This UV analysis corroborates the star-formation results from those of mid-IR to radio data for this very small sample.

References

Kaviraj, S. 2010, MNRAS, 408, 170

Leeuw, L. L., Davidson, J., Dowell, C. D., & Matthews, H. E. 2008, ApJ, 677, 249

Leeuw, L. L., Bregman, J., Davidson, J., Temi, P., & Im, S. S. 2011, in: C. Carignan, F. Combes & K. C. Freeman (eds.), *Tracing the Ancestry of Galaxies (on the land of our ancestors)*, Proceedings of the International Astronomical Union, *IAU Symposium*, Volume 277, p. 30