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ABSTRACT. An ICR investigation of the association reaction $CH_3^+ + HCN - CH_3 \cdot HCN^+$

has shown the reaction follows second order kinetics over the pressure range 1 x 10^{-6} to 3 x 10^{-4} Torr with a rate coefficient of 2 x 10^{-10} cm³s⁻¹. These results can be interpreted in terms of a saturated 3-body or radiative association mechanism.

Investigations of the reactions of the gaseous ion ${\rm CH_3}^+$ with a number of different molecules have shown the interesting feature that in several reactions the formation of the association product follows second-order kinetics¹,². The two main classes of association reactions that exhibit second order kinetics are (a) saturated 3-body reactions and (b) radiative association reactions. The application of radiative association reactions to interstellar processes provides a convenient means of forming large polyatomic molecules. However, even though radiative association has been invoked in models of interstellar chemistry³, there is still little direct experimental evidence for its participation as an important process. The 3-body association process represented by reaction (1) and radiative association by reaction (2)

$$A^{+} + B + M \xrightarrow{k_{1}} AB^{+} + M \tag{1}$$

$$A^{+} + B \xrightarrow{k_2} AB^{+} + hv \tag{2}$$

can be understood in terms of two distinct steps. The first of these steps (reaction (3)) is common to both 3-body and radiative association.

$$A^{+} + B \xrightarrow{k_{3}} (AB^{+})^{*}$$
 (3)

$$(AB^{+})^{*} + M \xrightarrow{k_{4}} AB^{+} + M$$
 (4)

$$(AB^{+})^{*} \xrightarrow{k_{5}} AB^{+} + h\nu \tag{5}$$

By using the steady state approximation applied to complex $(AB^+)^*$ and assuming that dissociation of $(AB^+)^*$ is rapid compared with stabilization, it is readily shown that the rate of production of the associated product

305

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306 M. J. McEWAN ET AL.

AB is given by

$$d[AB^{+}]/dt = (k_3/k_{-3})([A^{+}][B])(k_4[M] + k_5).$$
 (6)

The overall rate coefficients for 3-body and radiative association are therefore

$$k_1 = k_3 k_4 / k_{-3} \tag{7}$$

and

$$k_2 = k_3 k_5 / k_{-3}$$
 (8)

As typical values for the rate coefficients k_4 and k_5 are 10^{-9} cm³ s⁻¹ and 10^2-10^3 s⁻¹ respectively^{1,5}, and [M] $\sim 10^{15}$ molecules cm⁻³ in most flow tube experiments^{1,2}, then 3-body association should dominate radiative association.

We have applied the ion cyclotron resonance (ICR) technique to investigate the association process

$$CH_3^+ + HCN \longrightarrow CH_3 \cdot HCN^+$$
 (9)

and have found reaction (9) to follow second order kinetics throughout the entire pressure-range of the experiment, viz. $1 \times 10^{-6} - 2 \times 10^{-4}$ Torr. Good agreement was found between rate coefficients determined in the trapping mode at the low end of the pressure range and the drift mode at the high end of the pressure range. The observed rate coefficient at room temperature (300 K) was $k_9 = 2 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$ and was independent of pressure. A most interesting feature of this work is that the observed rate coefficient for k_9 is less than the value of 2 x $10^{-9} \, \mathrm{cm}^3 \, \mathrm{s}^{-1}$ reported by Schiff and Bohme² in their study of the much higher pressure range ~ 0.3 Torr. There are two likely explanations for the difference. In this work the product $k_4[M]$ is a factor ${\sim}10^4$ less than in the flow tube of Schiff and Bohme. Consequently we may be in a pressure regime where radiative association is the dominant mechanism for association whereas 3-body association is the major process at ~ 0.3 Torr. Alternatively, as the CH₃⁺ ions in our work were formed by the impact of low energy electrons on CH4, they may possess some vibrational excitation. A larger fraction of collisions between ${\rm CH_3}^+$ and HCN may then revert back to primary reactants (process k_{-3}) than was the case for thermal ions in the study by Schiff and Bohme.

If the second explanation is true, then the fact that we have observed second-order kinetics implies a value for τ_D , the lifetime of the complex (CH₃.HCN⁺)* to unimolecular decomposition at room temperature, of $\stackrel{>}{\sim} 3 \times 10^{-3}$ s.

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