

COSMOLOGICAL HII REGIONS

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We have generalized the classical description of ionization front propagation to the case of a point source in a uniform, cosmologically expanding gas. We present illustrative curves for the comoving radius and peculiar velocity for several turn-on redshifts, z_{ON} , for $\Omega_{tot} = 1$, $\Omega_b = 0.1$, $h = 1$. The quantity RS is the generalized Strömgen radius [$RS = RS_i (1 + z_{ON}) / (1 + z)$, $RS_i = (3N_u / 4\pi n_{H,i}^2 \alpha_2)^{1/3}$, $N_u =$ photoionizing number flux per source, $\alpha_2 =$ recombination rate to $n = 2$, $n_{H,i} = n_H^o (1 + z_{ON})^3$]. The quantity $T_{ON} = 2(1 + z_{ON})^{-3/2} / (3H_0)$. We also plot ζ , the value of $(2n_{Q,ph,Q}^o / 3H_0 n_H^o)$ needed to ionize the IGM with overlapping QSO HII regions by redshift z_{Ov} for QSO turn-on at various z_{ON} , where $N_{ph,Q}$ is ionizing photon luminosity per QSO, $n_Q^o =$ QSO number density (present co-moving value), $n_H =$ H density of IGM, and $n_H / n_H^o = n_Q / n_Q^o = (1 + z)^3$. From a recent preprint by Koo (1985), we estimate $\zeta \lesssim 1$ (for $\Omega_b = 0.1$, $h = 1$) for QSO's with $L \sim 10^{45}$ erg s⁻¹. In this case, the observed QSO's cannot be the sole source of the IGM ionization that is implied by the null detection of the Gunn-Peterson effect for QSO's with $z > 2$.

