

CCD ECHELLOGRAM OF THE STARBURST GALAXY TOL 1924-416

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The starburst galaxy Tol 1924-416 was observed on May 18, 1985 using an echelle spectrograph coupled with a cooled CCD camera (CASPEC) mounted at the Cassegrain focus of the ESO 3.6-m telescope. The echellogram covers the 4320-5260 Å spectral range with 0.33 Å/pixel and a spatial extent of 12 arcsec at 1.5 arcsec seeing. The estimated spectral resolution is 25 km s^{-1} and the velocity calibration error is as small as 3 km s^{-1} .

The obtained echellogram shows He I and He II emission in addition to strong Balmer lines and [O III] emission. Among others, a high optical depth line He I 5016 is detected.

A preliminary analysis has revealed that the mean heliocentric velocity is $2843 \pm 9 \text{ km s}^{-1}$. A small but distinct rotation of about $4 \text{ km s}^{-1} \text{ arcsec}^{-1}$ has been found for the central region. The FWHM of the emission lines is about 85 km s^{-1} . The emission line profiles are slightly asymmetric.

The line ratio of [O III] 5007/4363 gives $T_e = 13\,500 \text{ K}$. The line ratio of [O III] 5007/4959 is found to be 3.2 ± 0.1 , deviating from the theoretical value of 2.9.

The present high resolution spectroscopic observation thus provides spatially resolved kinematical and physical information for the first time on the nuclear H II region of the starburst galaxy Tol 1924-416. The full paper will be published elsewhere.