Luminosity-Distances of IUE observed Active Galaxies

Vijayakumar H Doddamani and P. Vedavathi

Dept. of Physics, Bangalore University, Bangalore-560056 email: drvkdmani@gmail.com

Abstract. Active galaxies are the most luminous objects observed in the Universe and are believed to be powered by mass accretion processes taking place in the vicinity of the central Super massive black hole ($M_{BH} \ge 10^8 \, M_{sun}$). However, the details of the power generation mechanisms are not understood well yet. In this paper, we are presenting a comparative study of luminosity-distance estimations for the complete sample of active galaxies observed by IUE satellite by different methods. IUE has made UV spectroscopic observations of nearly 400 active galaxies comprising mostly Seyfert 1 galaxies and quasars. We have chosen all the active galaxies observed by IUE satellite for the study of luminosity-distance with redshift. The luminosity-distances (D_L) have been calculated using the Hubbles law under non-relativistic and relativistic limits with $H_0 = 73 \, \text{Km/sec/Mpc}$ and Terrell (1979) also. We have found that all D_L estimations are consistent with each other for $z \le 1$ and diverge for $z \ge 1$. The results of cosmological calculator I and II are found to consistent with each other and higher by several factors over cosmological calculator IV and the predictions of the Hubble's law under relativistic case. We observe a kind bimodal distributions in D_L for $z \le 3.5$.

Keywords. Active galaxies, Seyfert 1 galaxies, quasars, low redshift, Luminosity-distance

1. Introduction

The active galaxies are the most luminous objects observed in the Universe and are believed to be powered by mass accretion processes taking place in the vicinity of the central super massive black hole ($M_{BH} \geqslant 10^8 M_{sun}$). Thus, it becomes very much important to understand the issue of their luminosity-distance (D_L) estimations. There are no direct methods available for the distance estimation of active galaxies independent of their redshifts. The redshift-independent distance estimation methods are reliable only for the galactic scale. Hubbles law cannot be used in its simple form to estimate D_L estimations at higher redshift active galaxies. Active galaxies serve as standard-distance candles for the distance estimations on cosmological scales. To look into the consistent D_L estimation issues of active galaxies especially for $z \geqslant 1$, we have undertaken the D_L estimations of the complete sample of IUE observed active galaxies using simple Hubbles law under both non-relativistic and relativistic limits (Bogess *et al.* 1978). We have made use of the following relations for D_L estimations:

$$D_L = cz/H_0 \tag{1.1}$$

$$D_L = V_R z / H_0 \tag{1.2}$$

$$(1+z)^2 = (c+V_R)/(c-V_R)$$
(1.3)

$$V_R = Kc \tag{1.4}$$

where K- a numerical constant and c-velocity of light. These estimations are compared with the results of D_L estimations using Terrell (1979) formula and cosmological

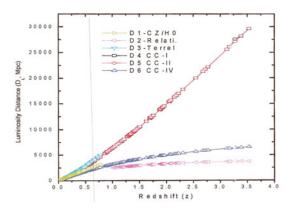


Figure 1. Dependence of Luminosity-Distances (DL, Mpc) with redshift (z) for IUE observed active galaxies.

calculators available online at http://ned.ipac.caltech.edu. A cosmological model of $\omega = 0.7$ and $\omega_m = 0.3$ with $H_0 = 73$ Km/sec/Mpc has been assumed for all the calculations.

2. IUE Database

The IUE satellite has made low resolution UV spectroscopic observations of Seyfert galaxies (Sy1 and Sy2), Quasars, BL Lac HP objects, Radio galaxies and HII galaxies in the 1200-3000 Å wavelength range (Bogess A. et al. 1978). In order to study the line luminosity dependence with the UV continuum, we have undertaken luminosity-distance D_L calculations for all the 412 active galaxies for their redshift $z \leq 3.5$.

3. Conclusions and Discussions

We have found that all D_L estimations are consistent with each other for $z \leq 0.6$ from and diverge for $z \geq 0.6$. The results of cosmological calculator I and II (D_4 and D_5 in fig. 1) are found to consistent for all values of z with each other. The D_L estimations from Hubbles law (D_1) under non-relativistic limit, cosmological calculator IV (D_6) and relativistic Hubble law (D_2) have all yielded lower values by higher by several factors over cosmological calculator I and II. Thus, we have found a kind of bimodal distributions in D_L for $z \leq 3.5$. The results imply that the D_L estimations methods are not able to produce unique distance value for an object in higher redshift regime ($z \geq 3.5$).

Acknowledgements

This work has been carried out under a UGC major research project sanctioned to the first author. We also acknowledge financial support of UGC, New Delhi, and Bangalore University with gratitude towards carry out this research work.

References

Bogess, A., et al. 1978, Nature 275, 372 & 377

Peterson, B. M. 1997, An Introduction to Active Galactic Nuclei (CUP)

Schmidt, B. & Riess, A. 1998, AJ 116, 1009

Cappi, A. 2001, ApL&C 40, 161

Perlmutter, S., et al. 1999, ApJ 517, 565

Perlmutter, S., et al. 1995, ApJ 440, L41

Terrell, J. 1977, A. J. P. 45, 869

Wright, E. L. 2006, PASP 118, 1

Wright, E. L. 2006, PASP 118, 1711

Doddamani, V. H. 2004, *Ultra Violet Spectroscopic study of Seyfert 1 galaxies and quasars* (Thesis submitted to Bangalore University), Bangalore, India